

The occurrence rate of hot Jupiters from TESS

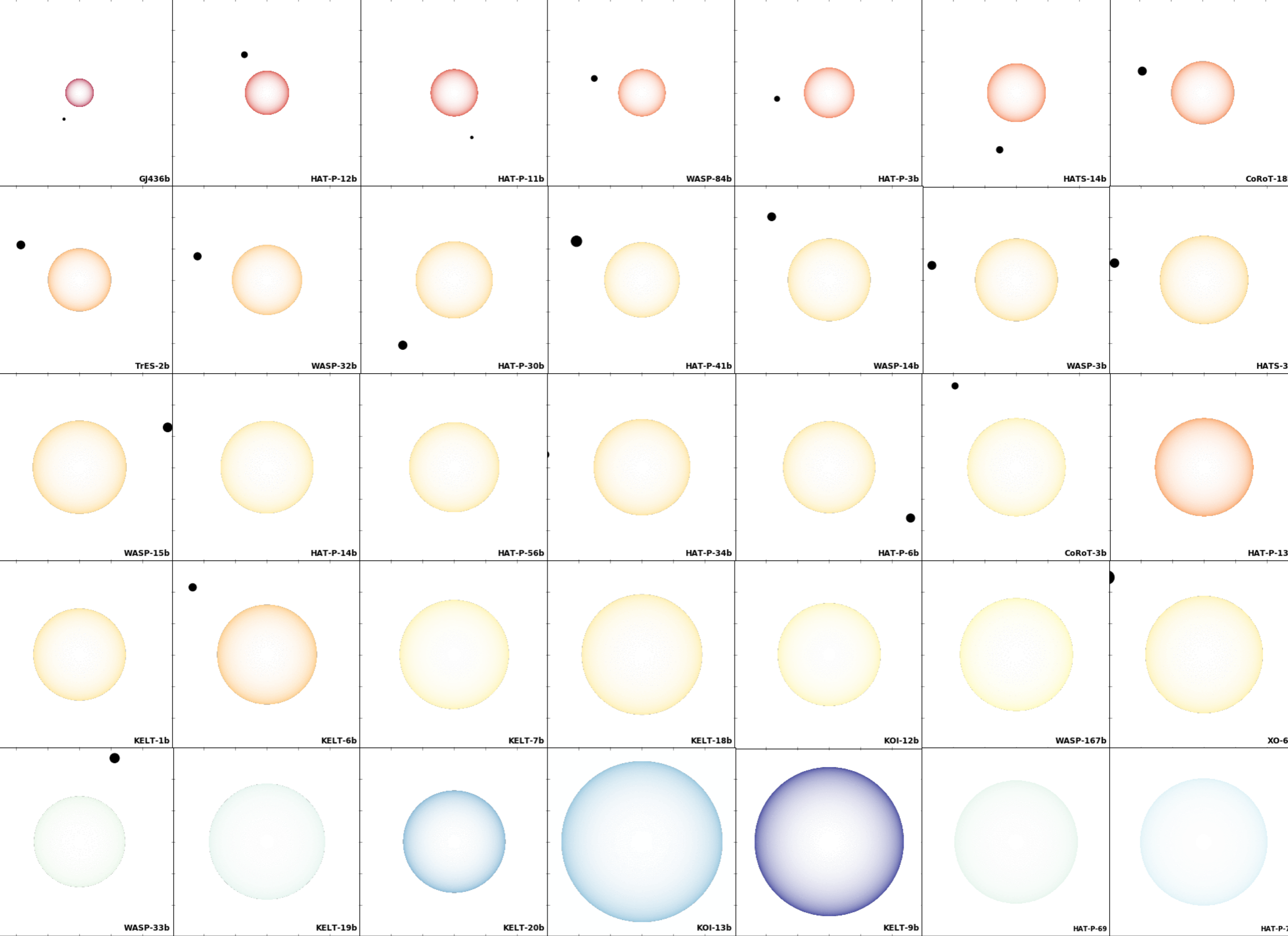
George Zhou

CENTER FOR

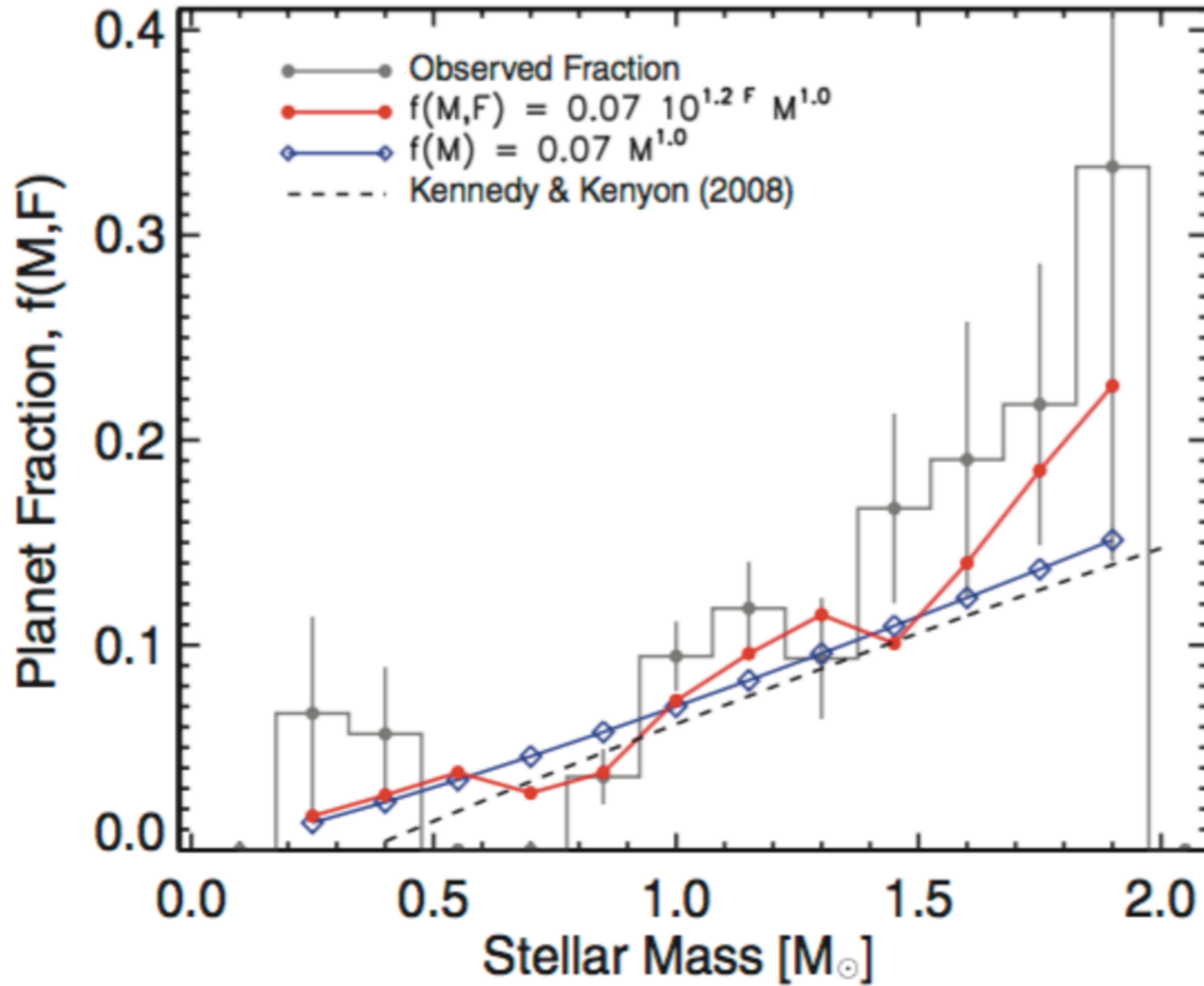
ASTROPHYSICS

HARVARD & SMITHSONIAN

Hubble Fellow



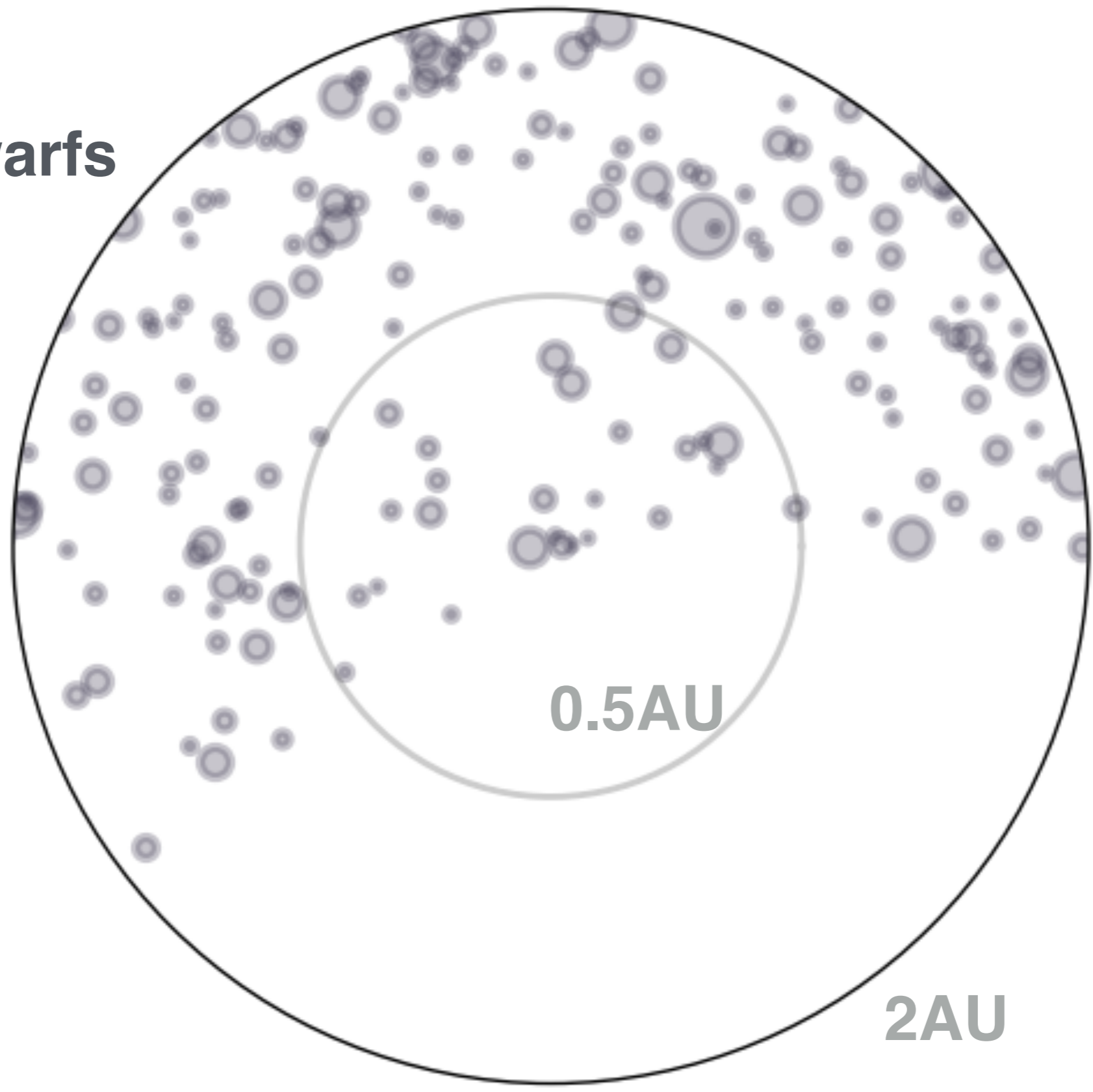
Retired A-stars have more giant planets



Johnson+ 2010, also see Jones+ 2014

Giant planets from RV

Dwarfs

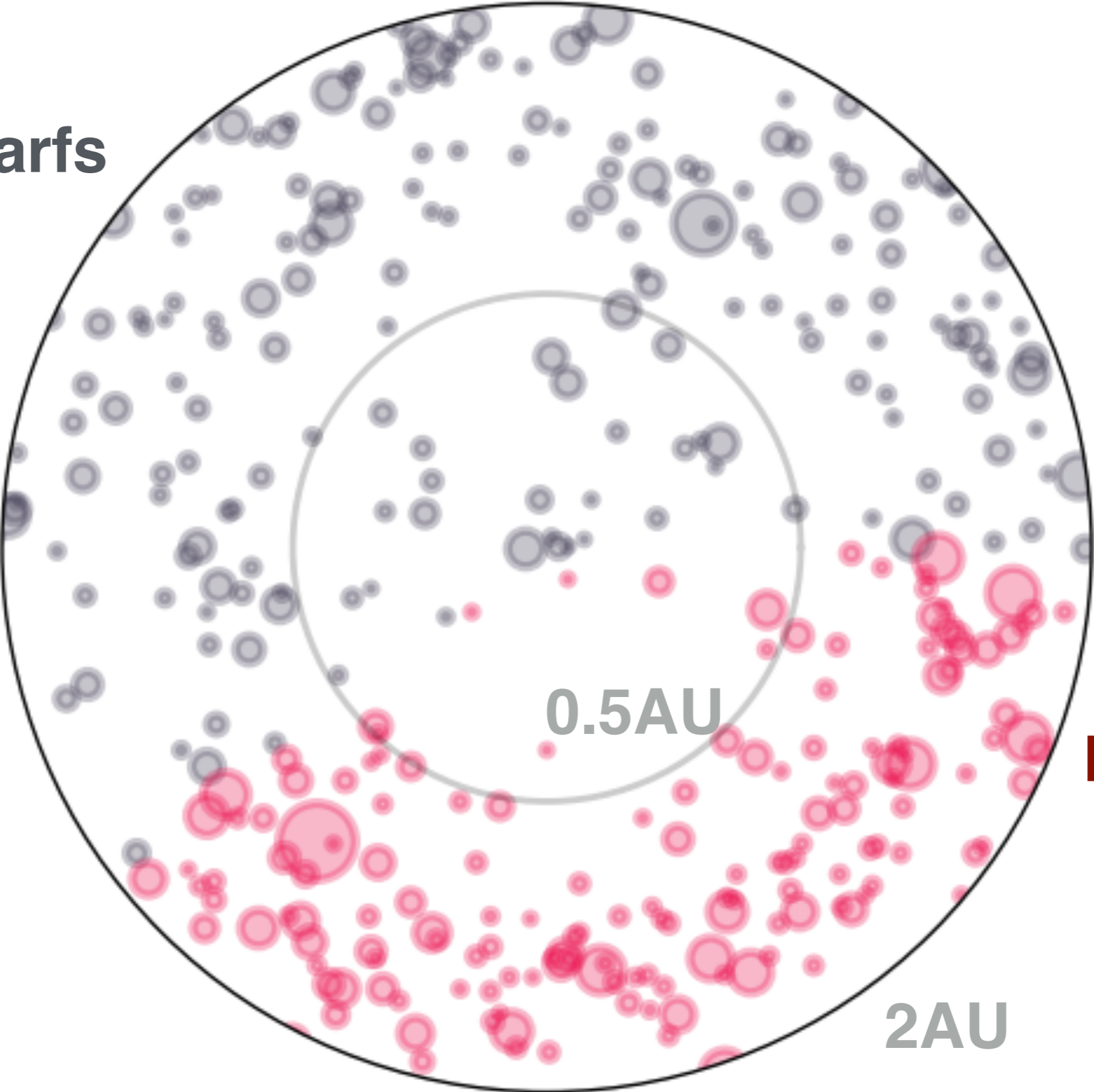


0.5AU

2AU

Giant planets from RV

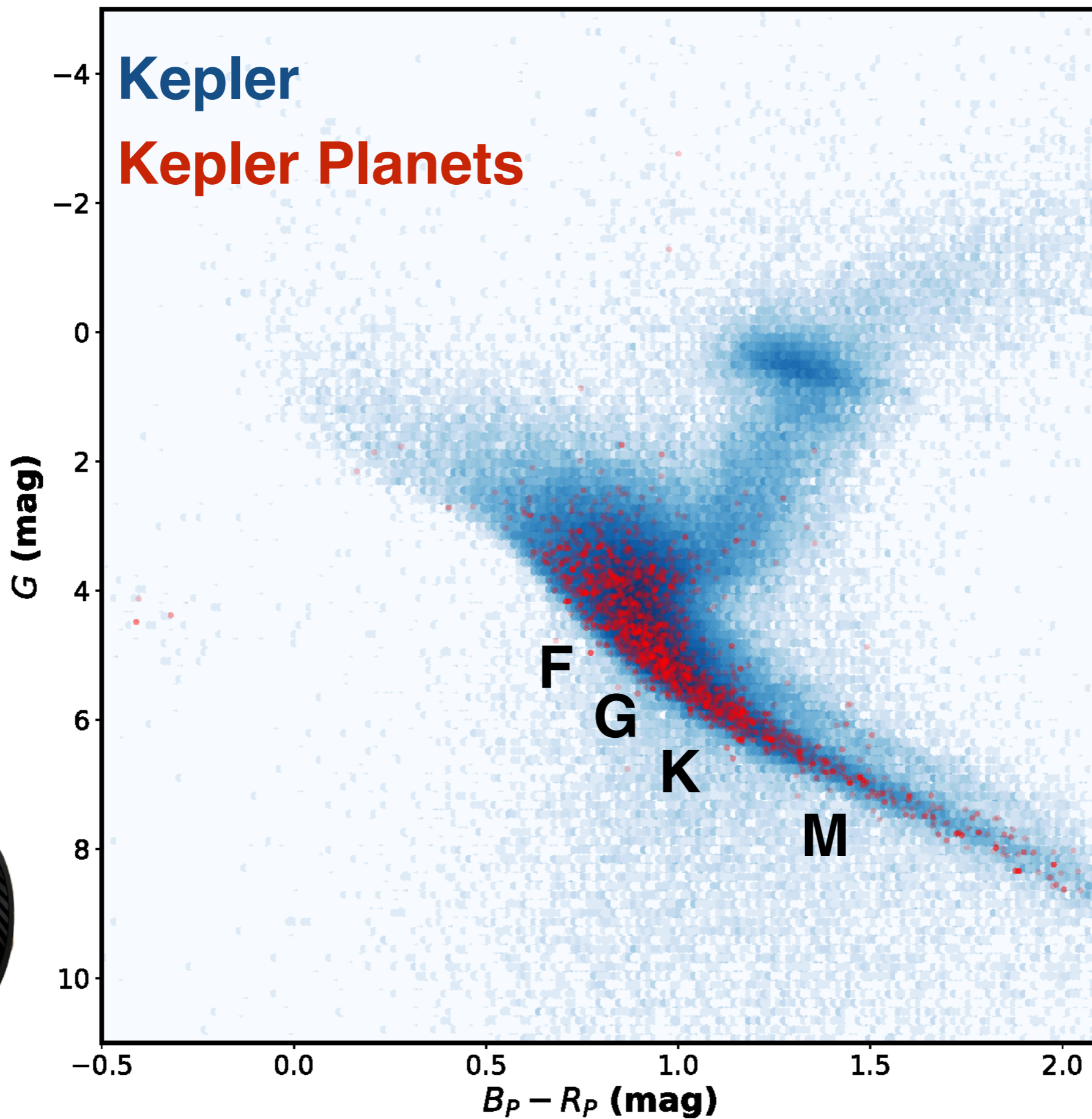
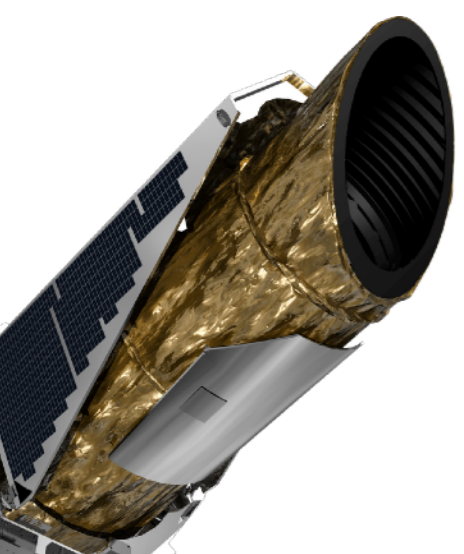
Dwarfs

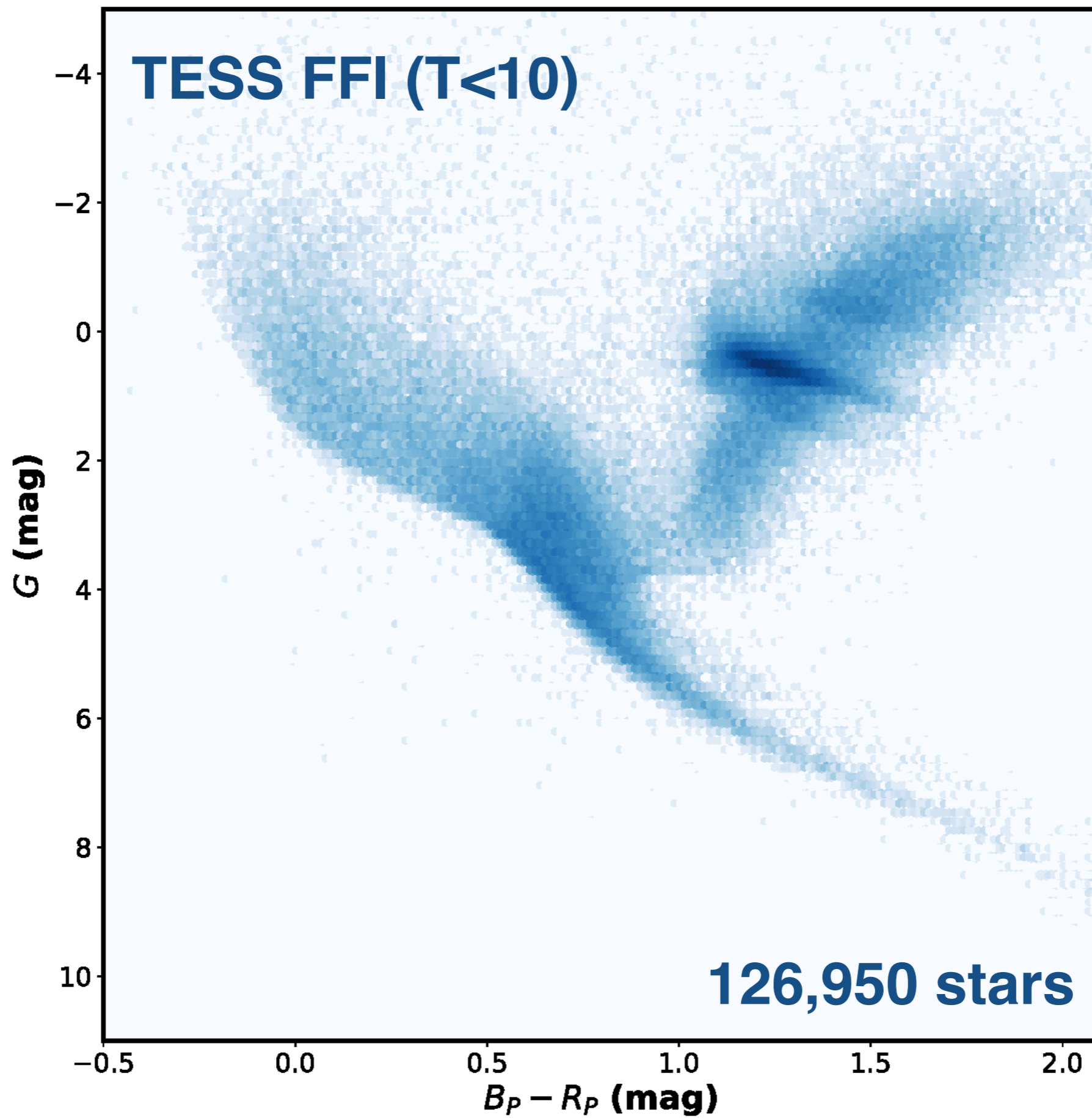
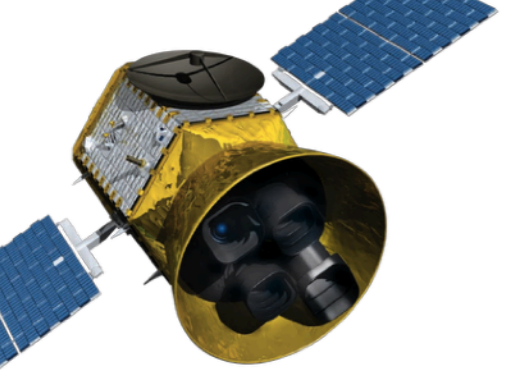


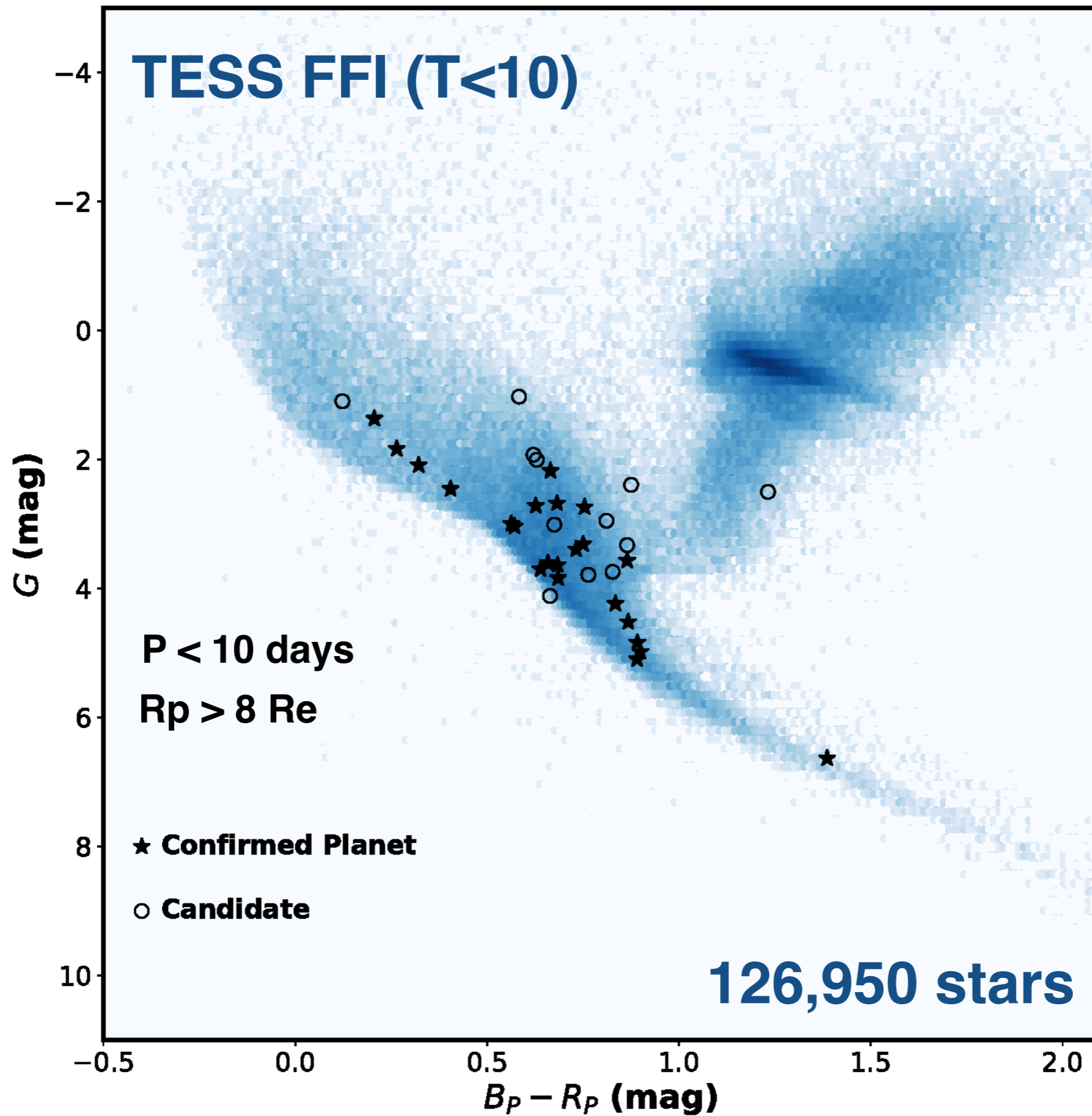
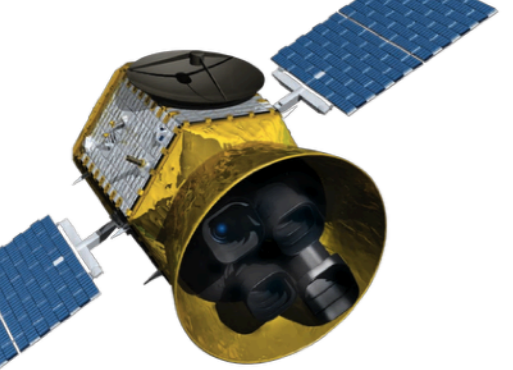
0.5AU

Evolved

2AU





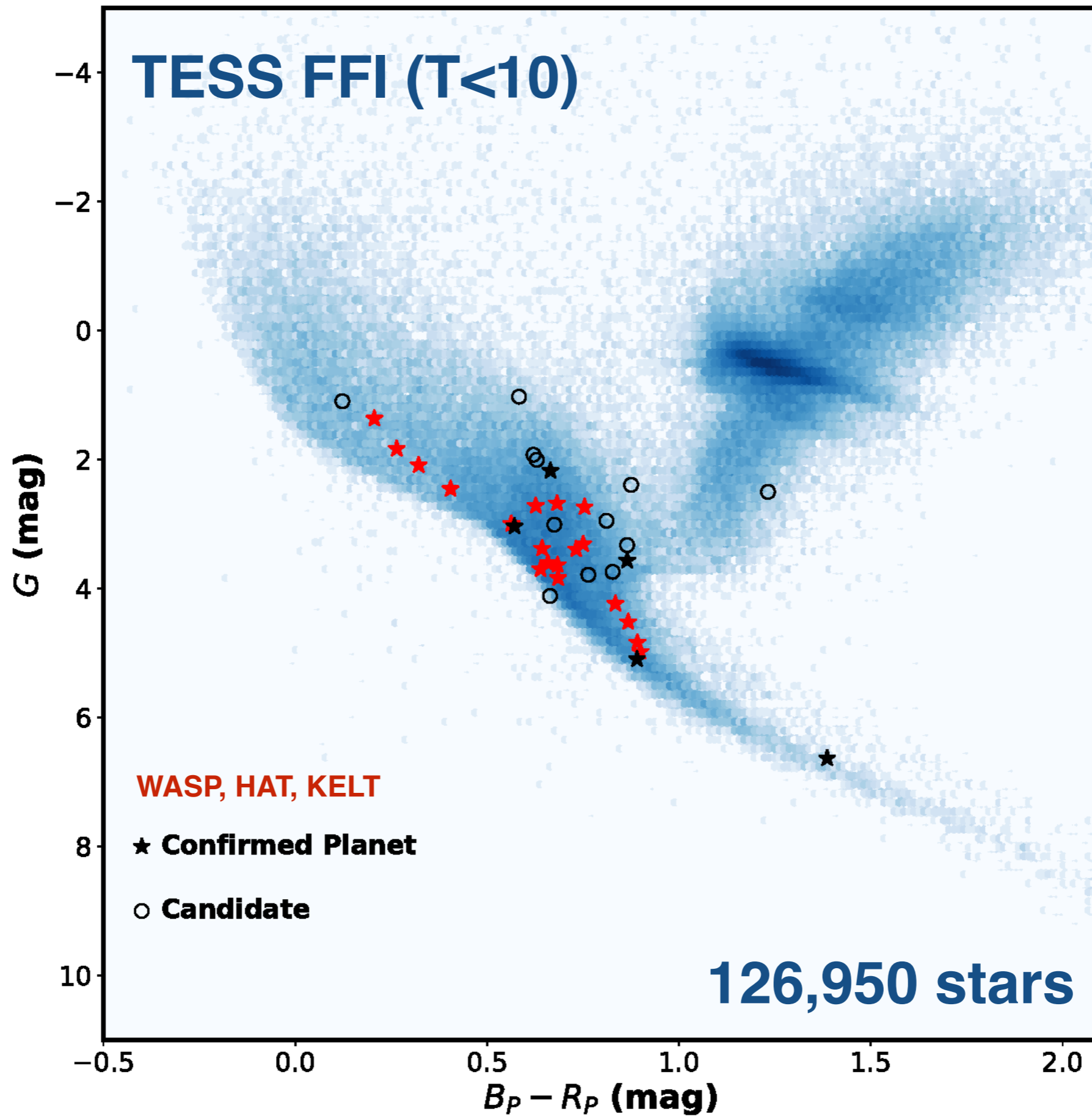
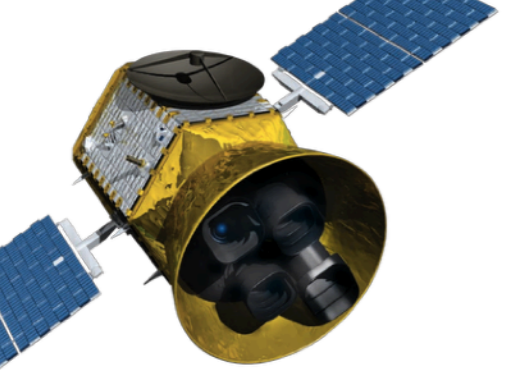


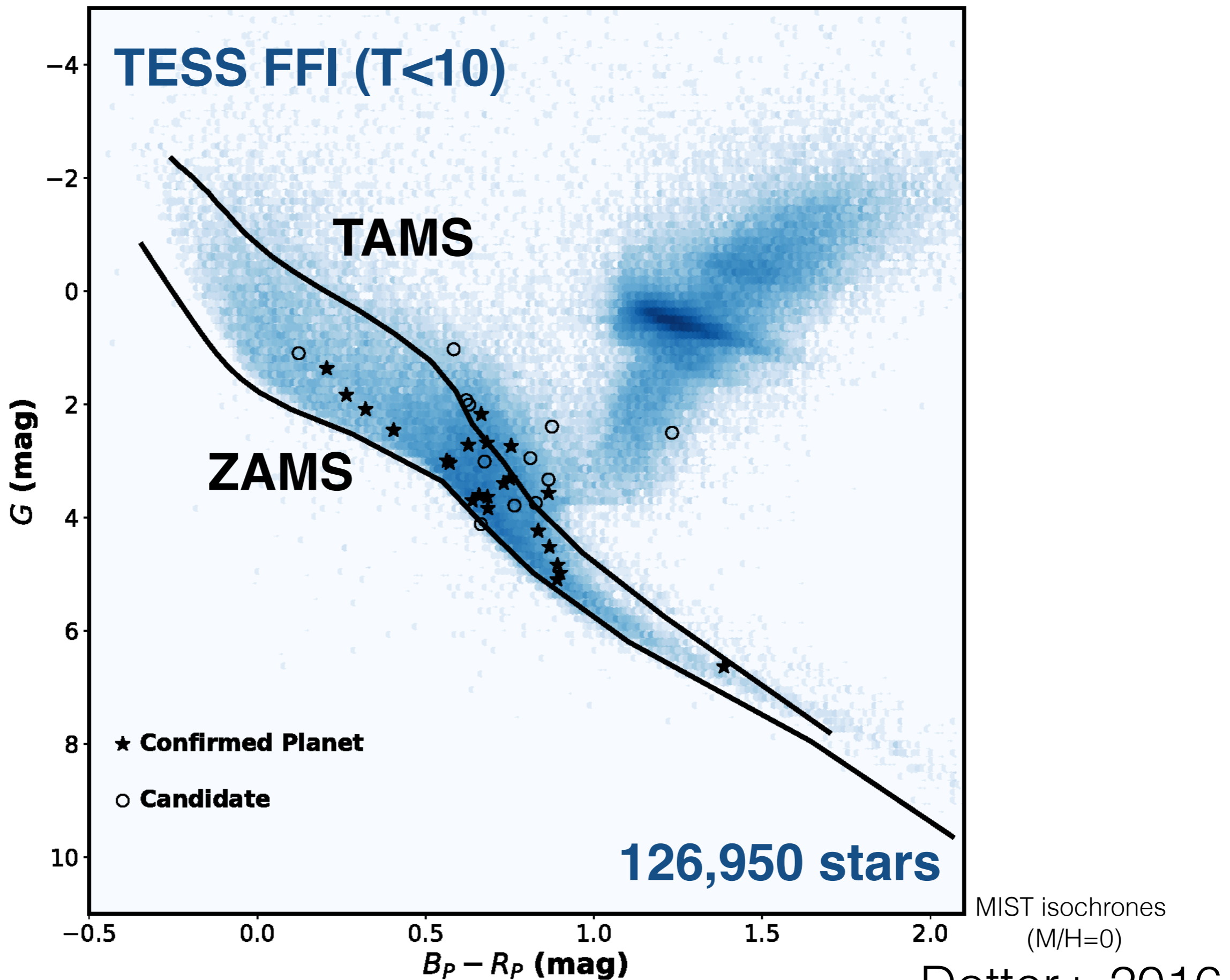
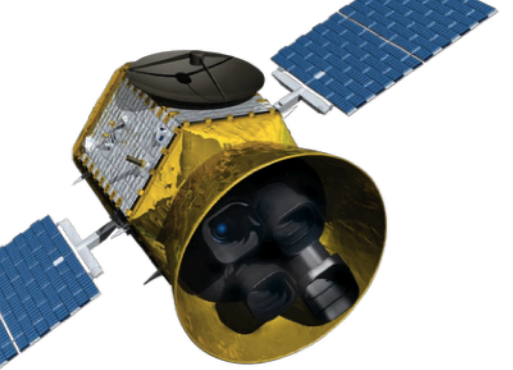
31 candidates

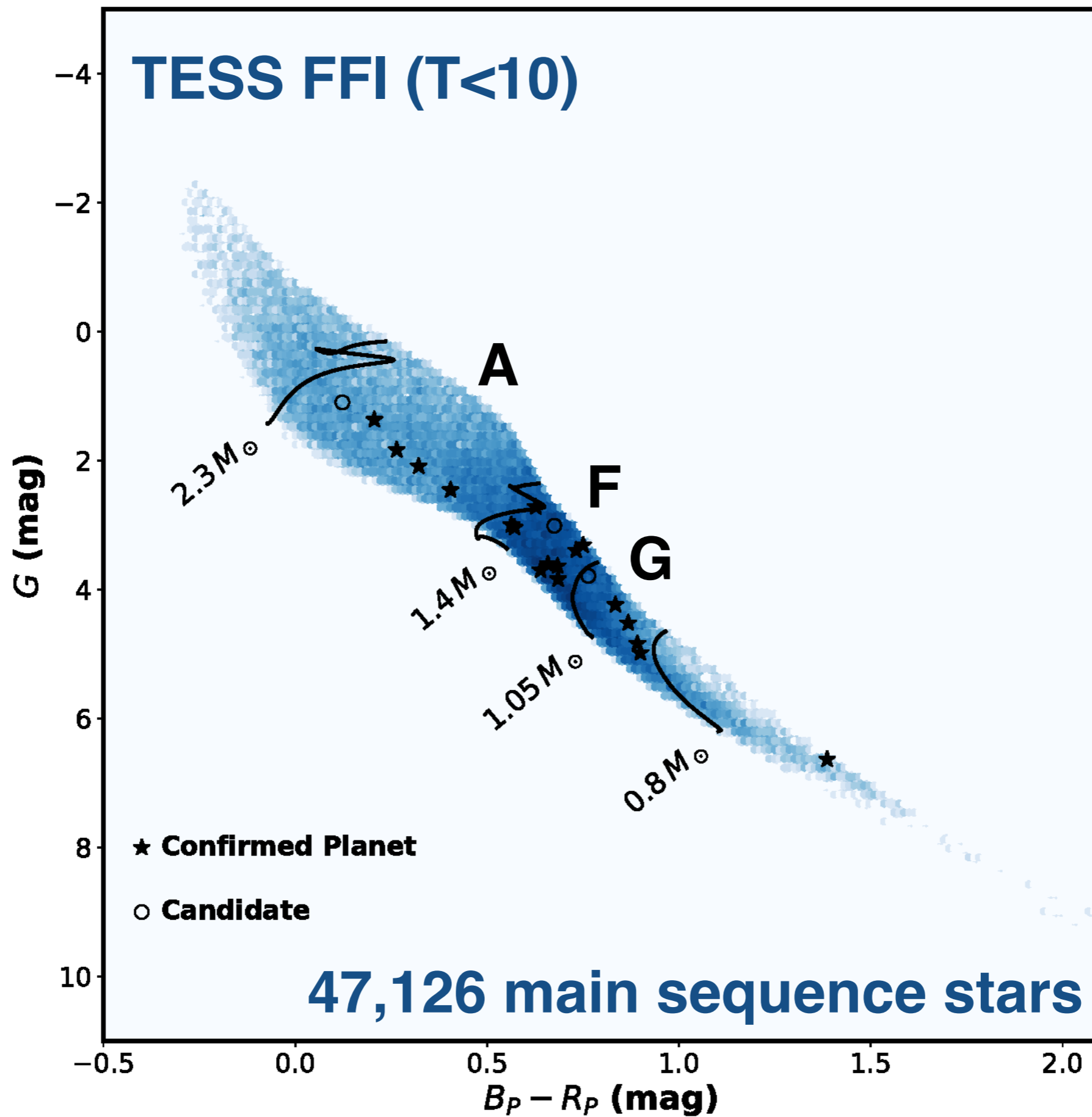
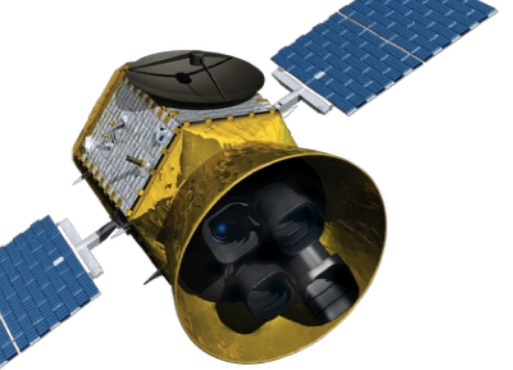
18 confirmed

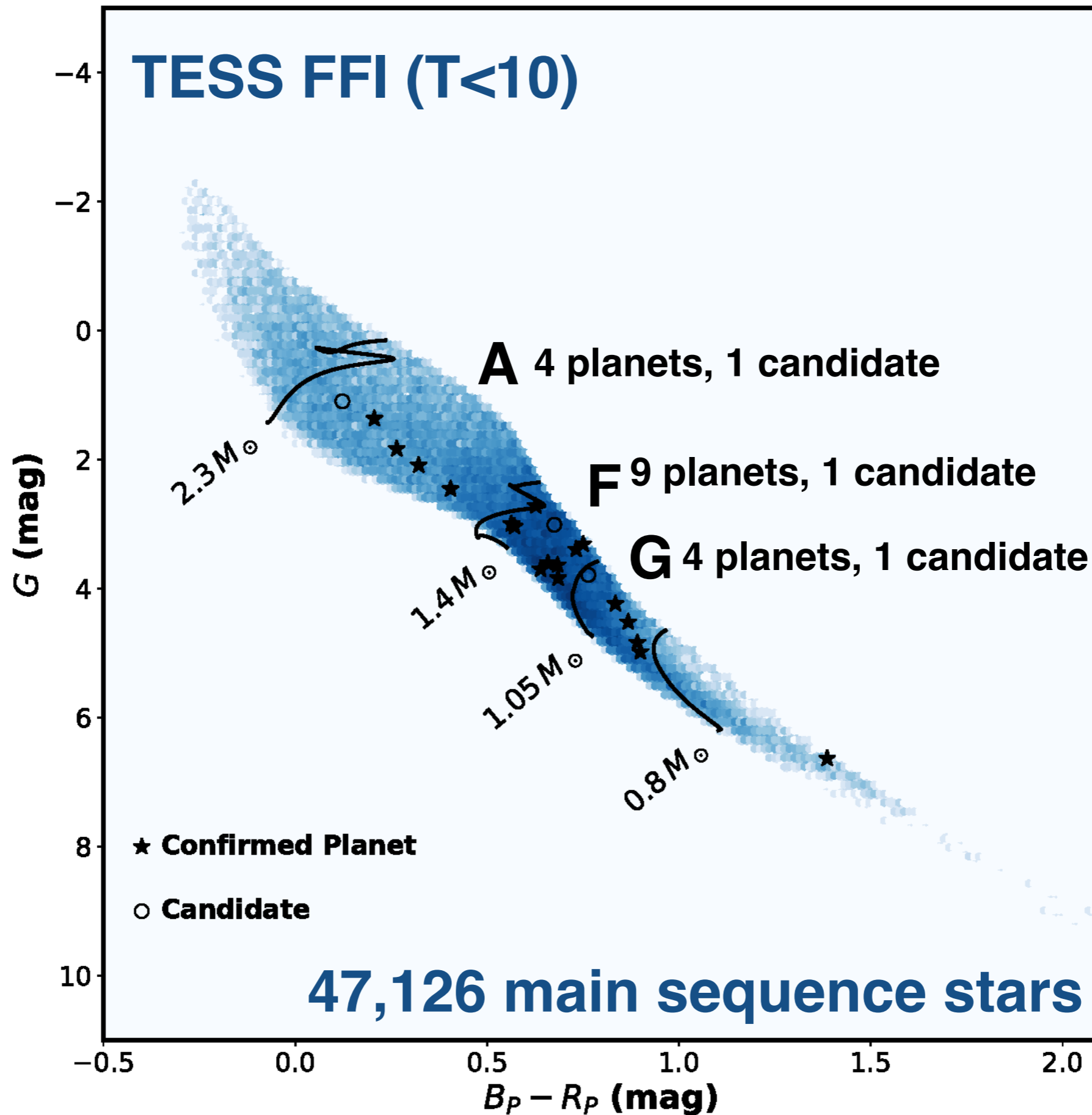
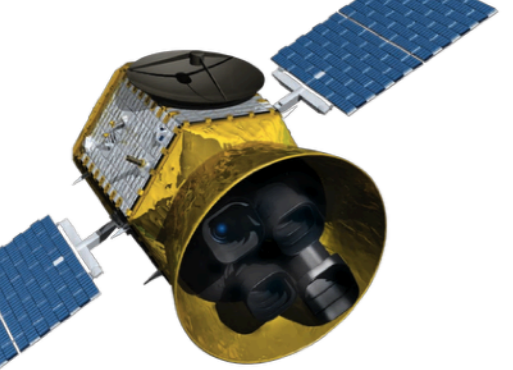
3 active

**10 False
Positives**



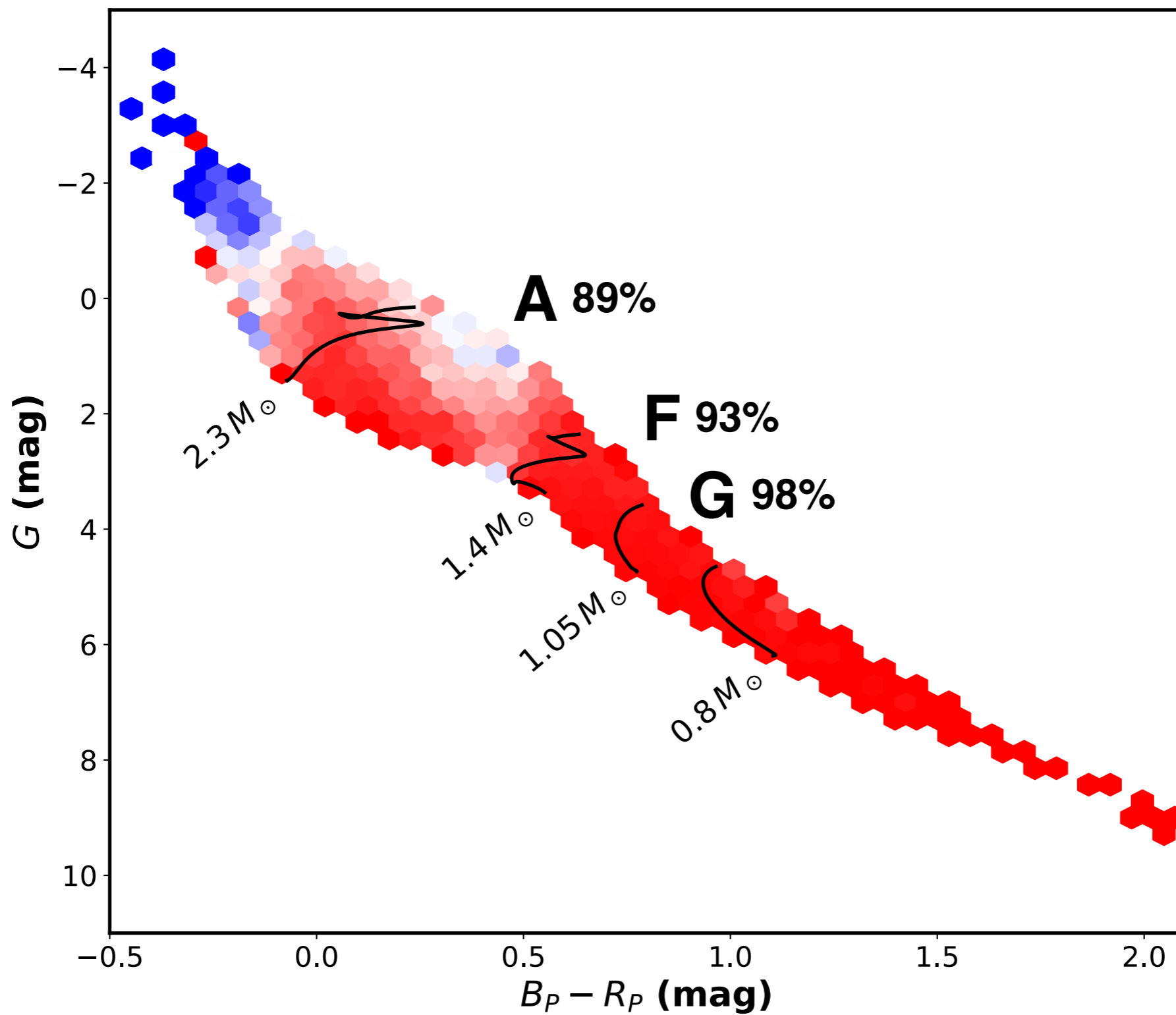




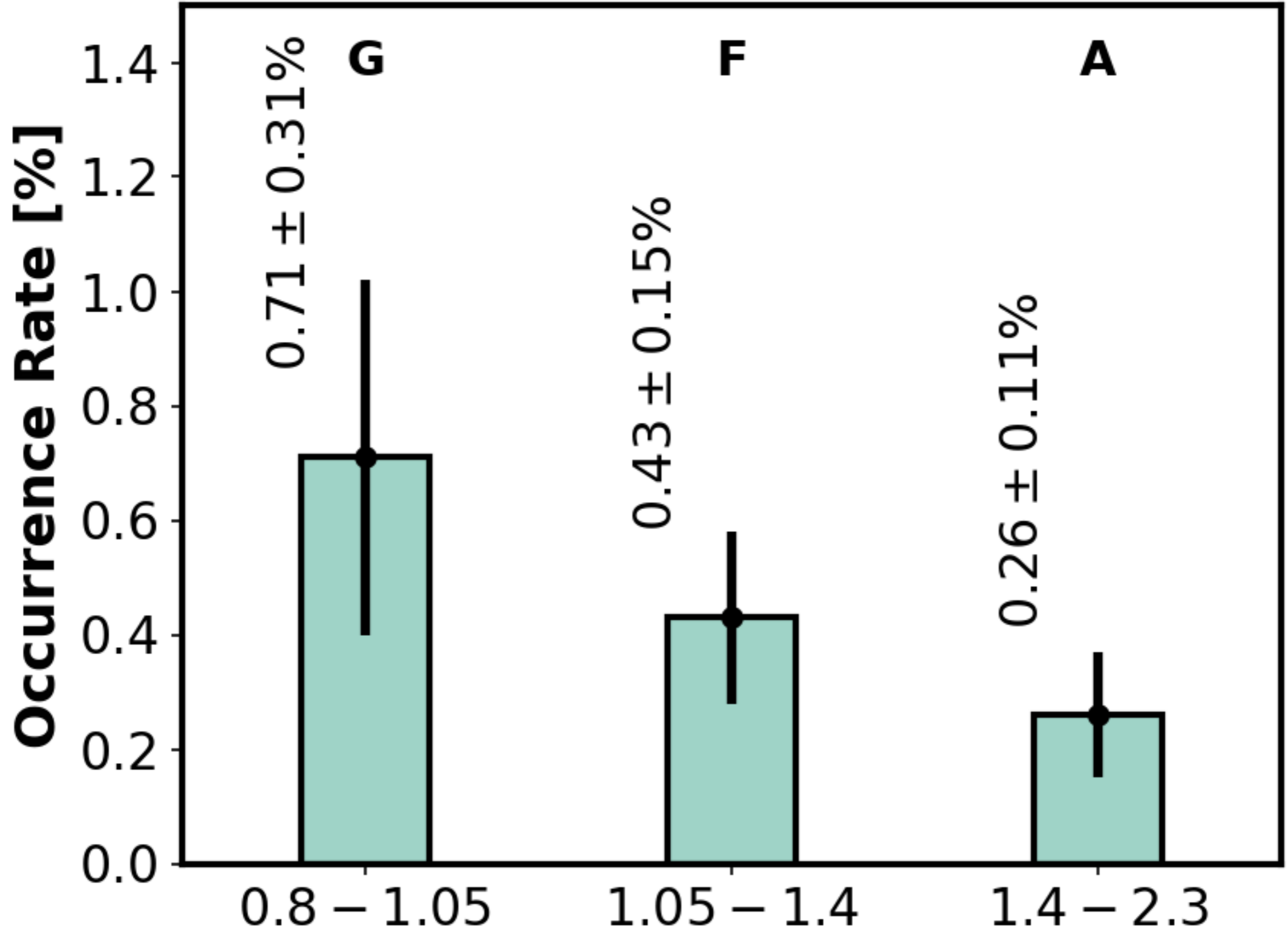


Completeness of 10 day Jupiters

0.0 0.5 1.0

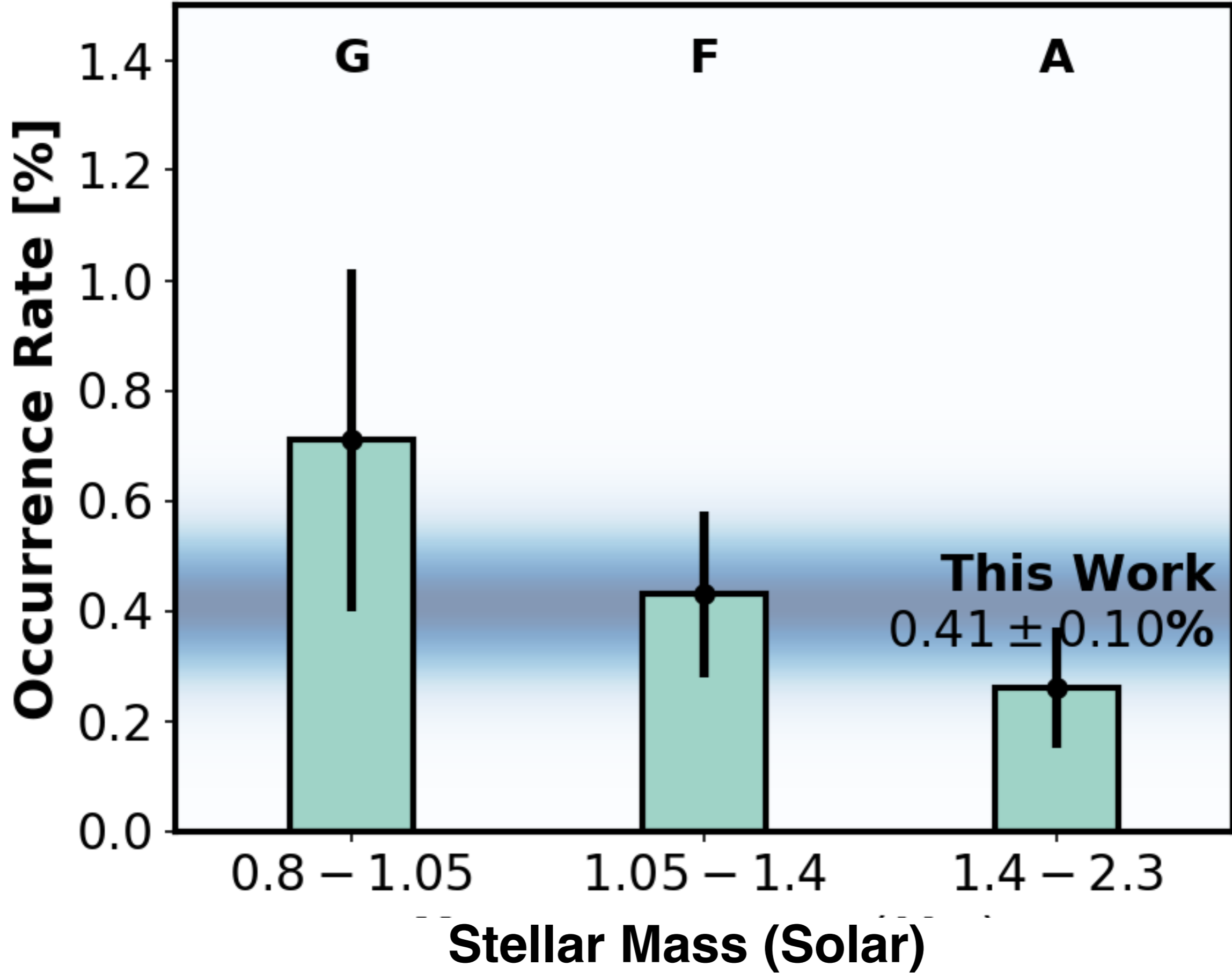


TESS HJ Occurrence Rate

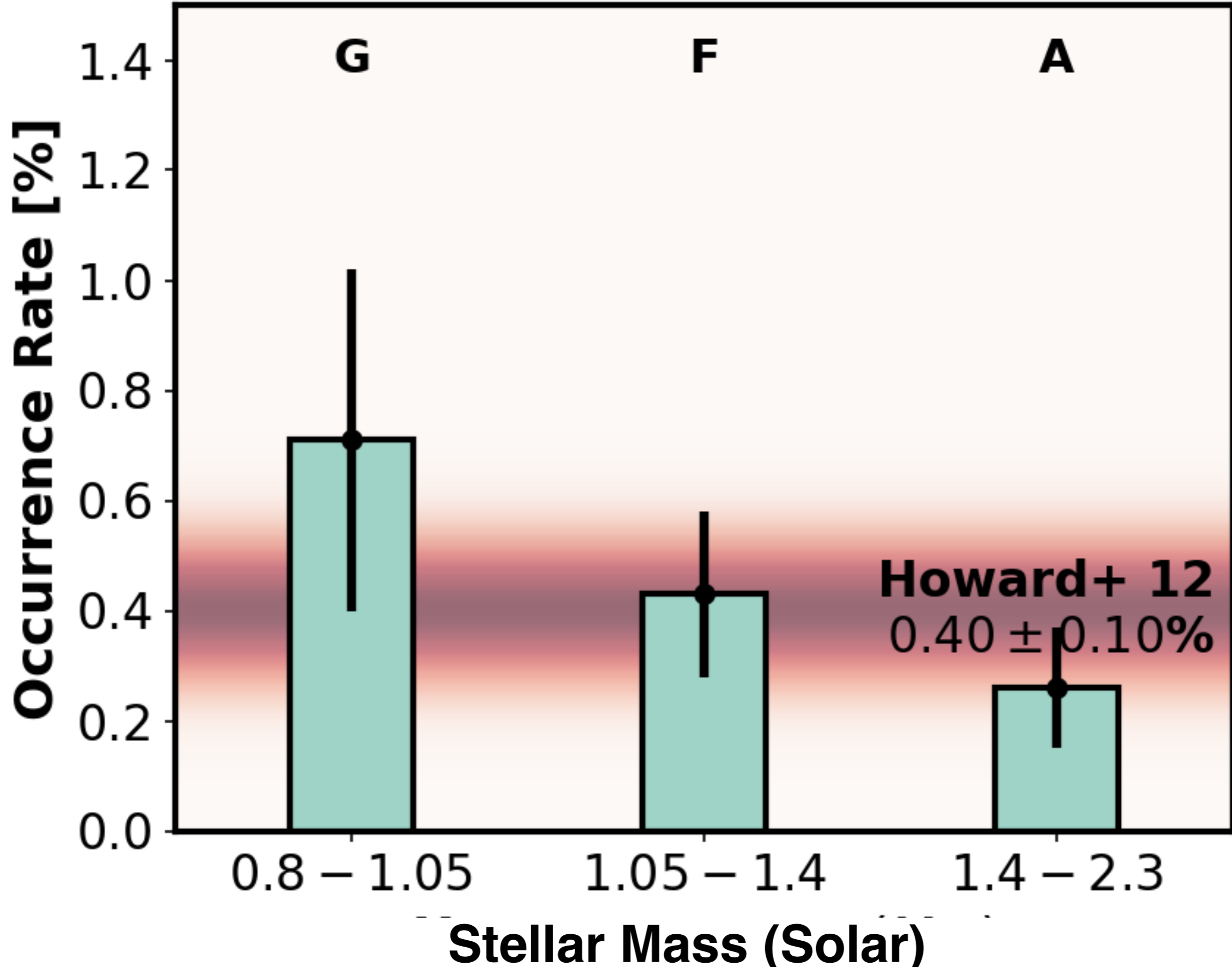


Stellar Mass (Solar) Zhou et al. (2019)

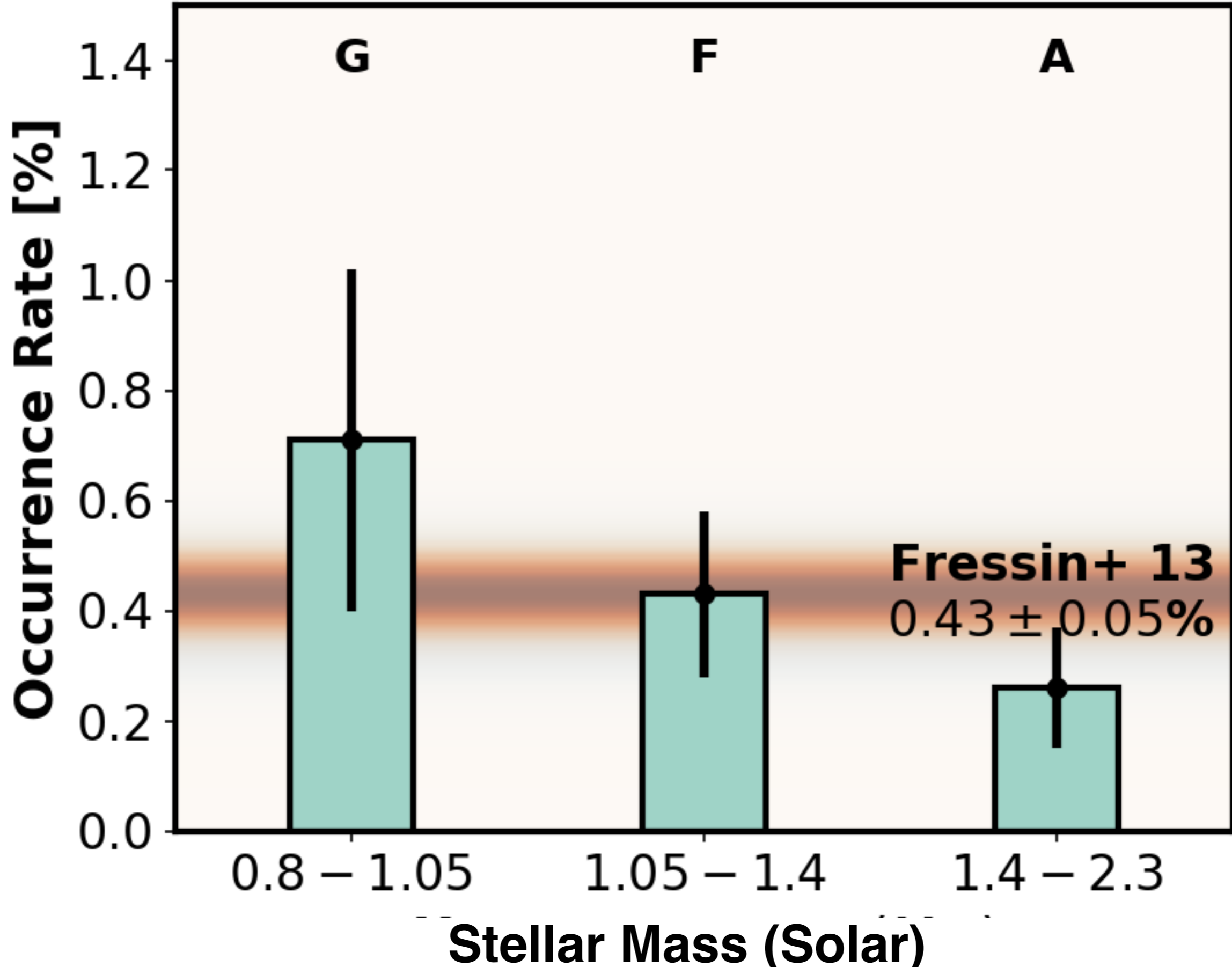
TESS HJ Occurrence Rate



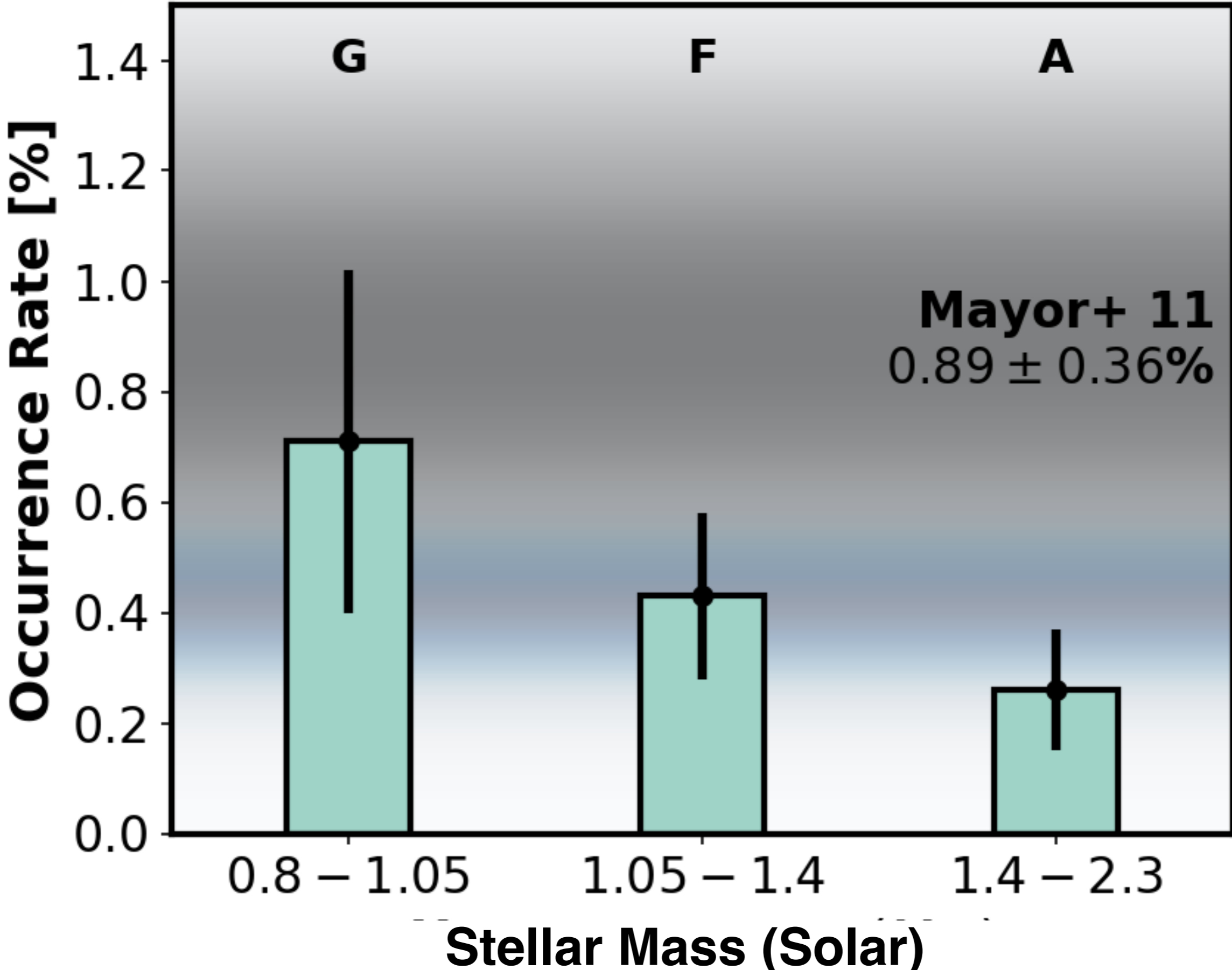
TESS vs Kepler HJ Occurrence Rate



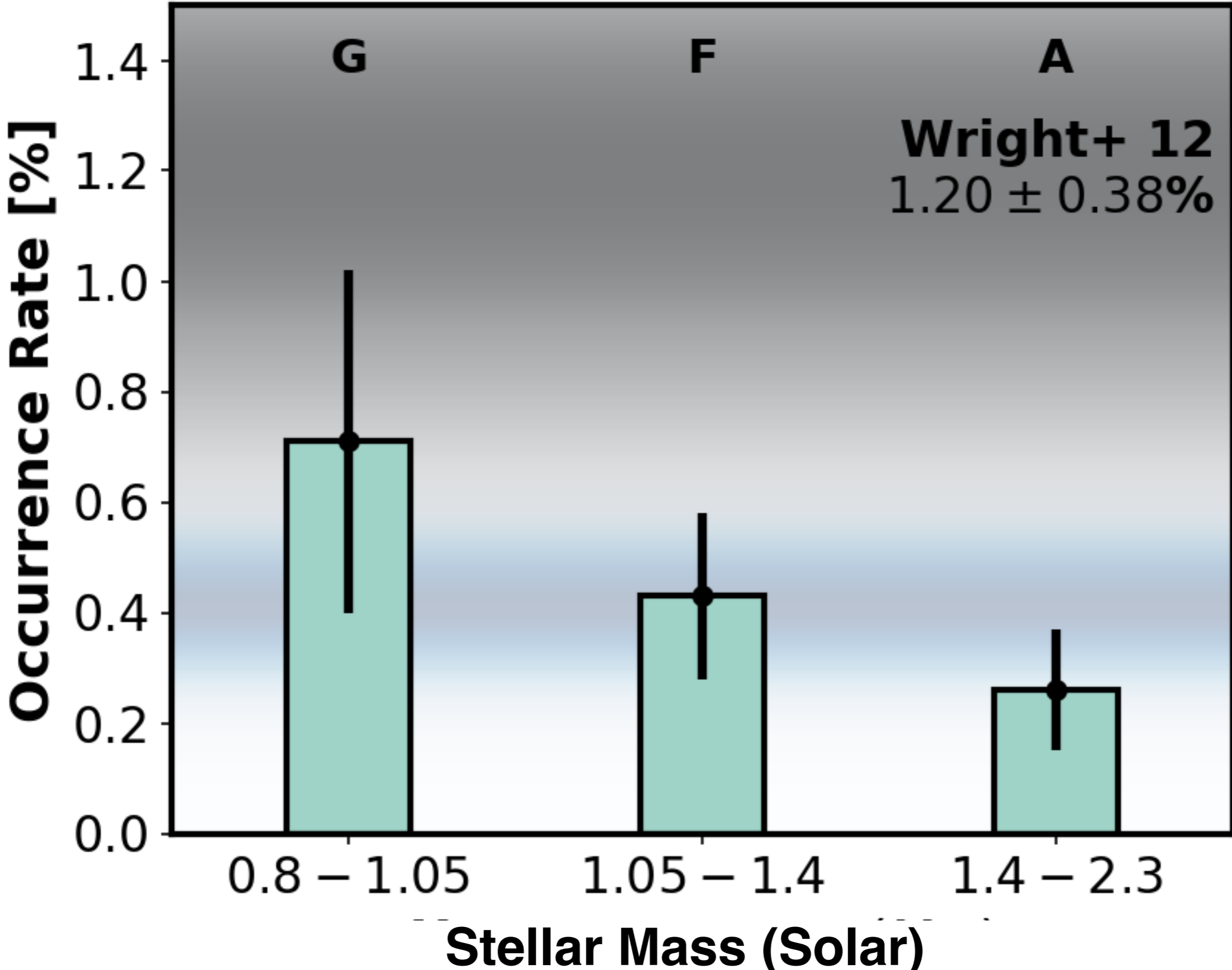
TESS vs Kepler HJ Occurrence Rate

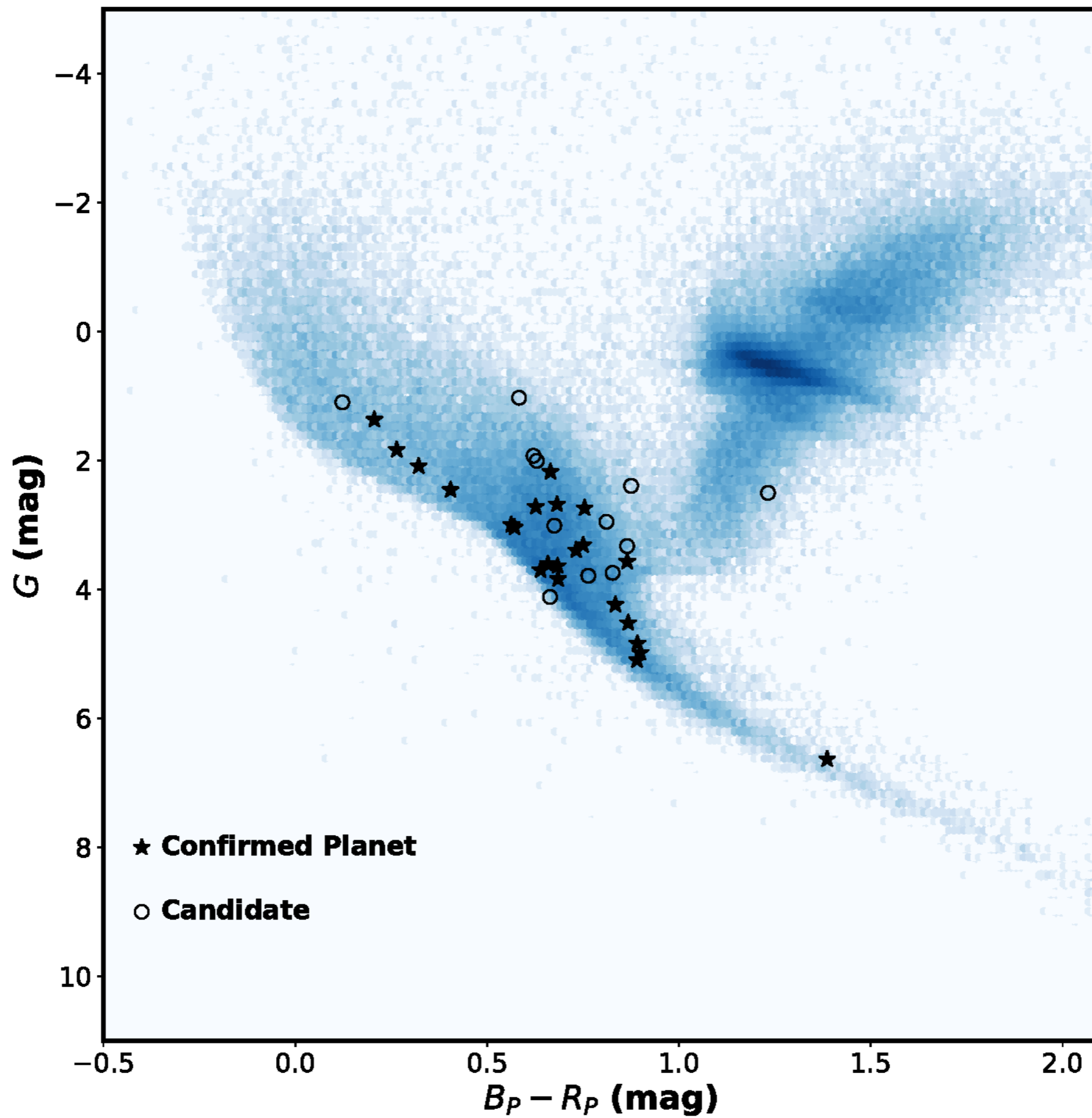


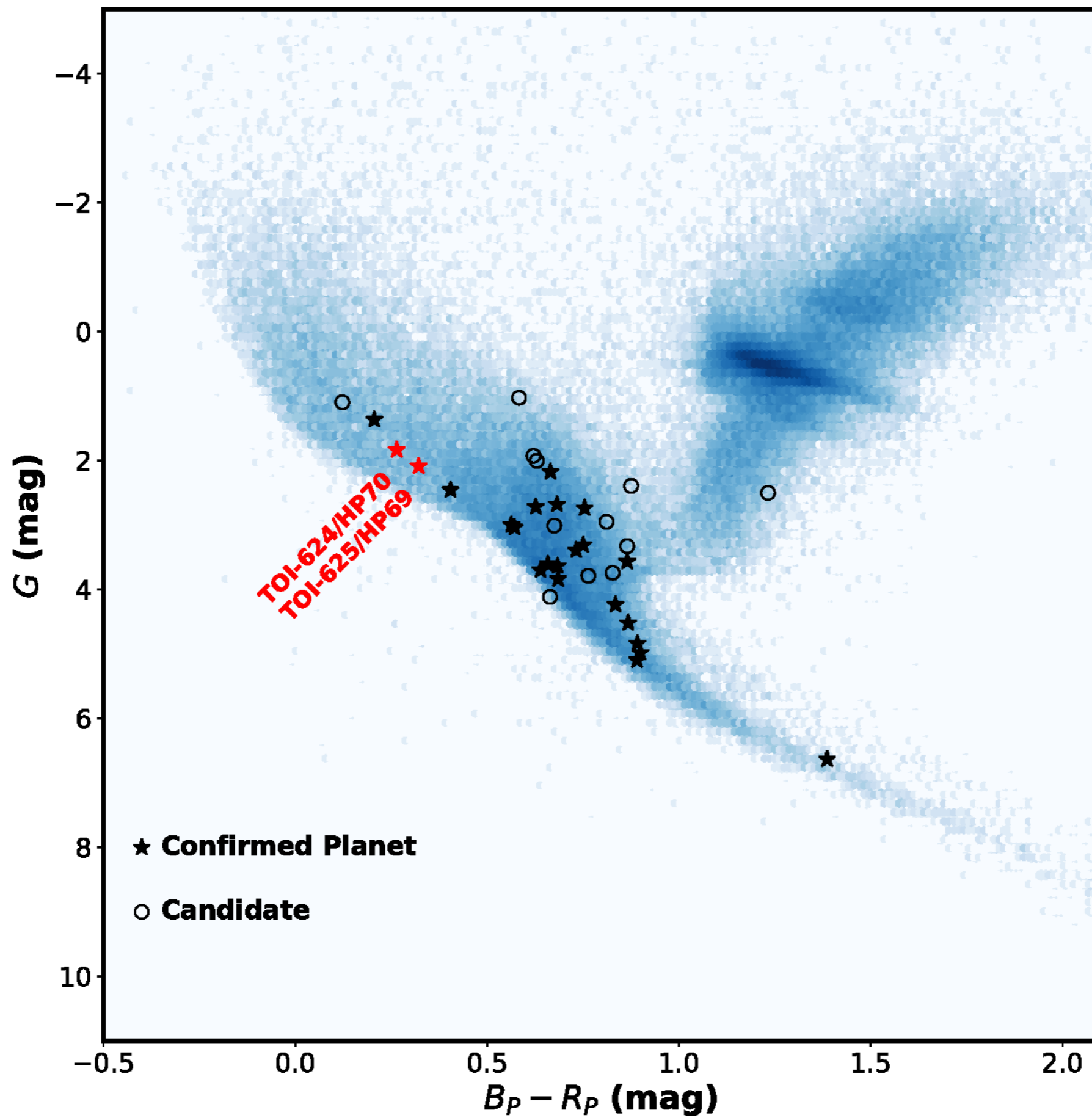
TESS vs RV HJ Occurrence Rate

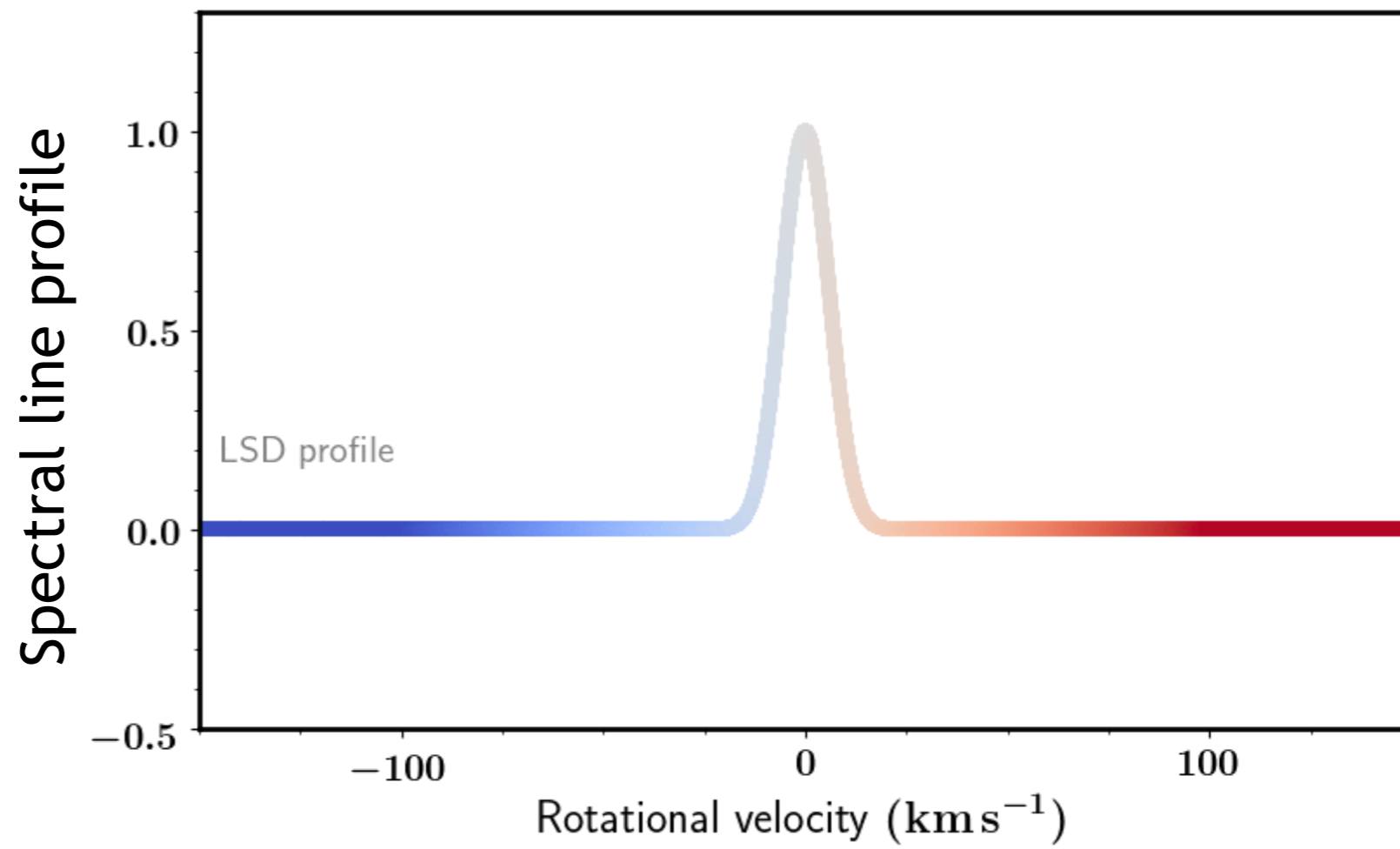


TESS vs RV HJ Occurrence Rate

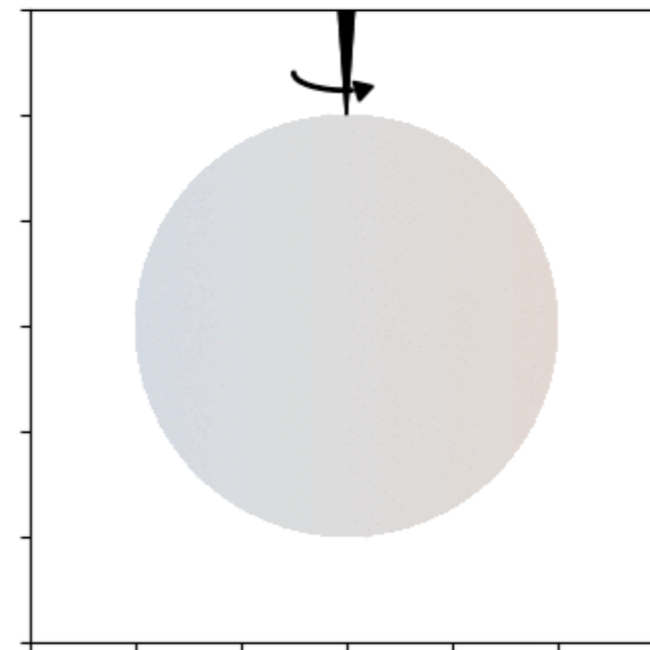




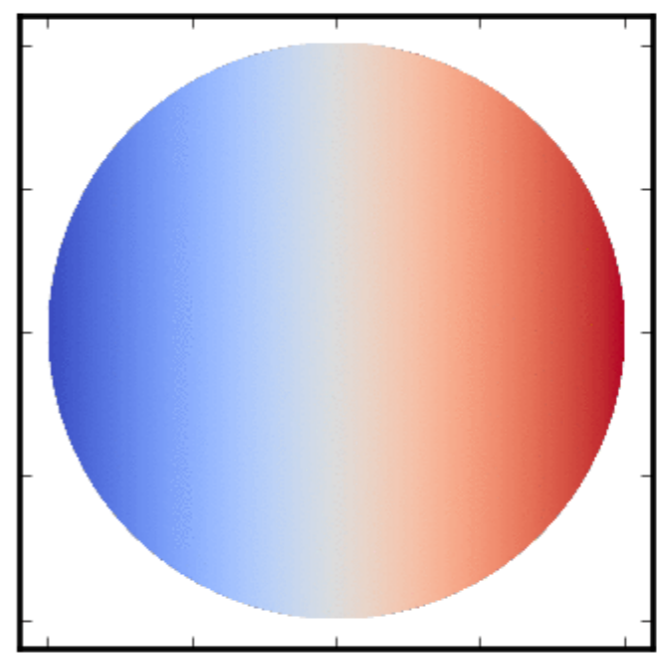
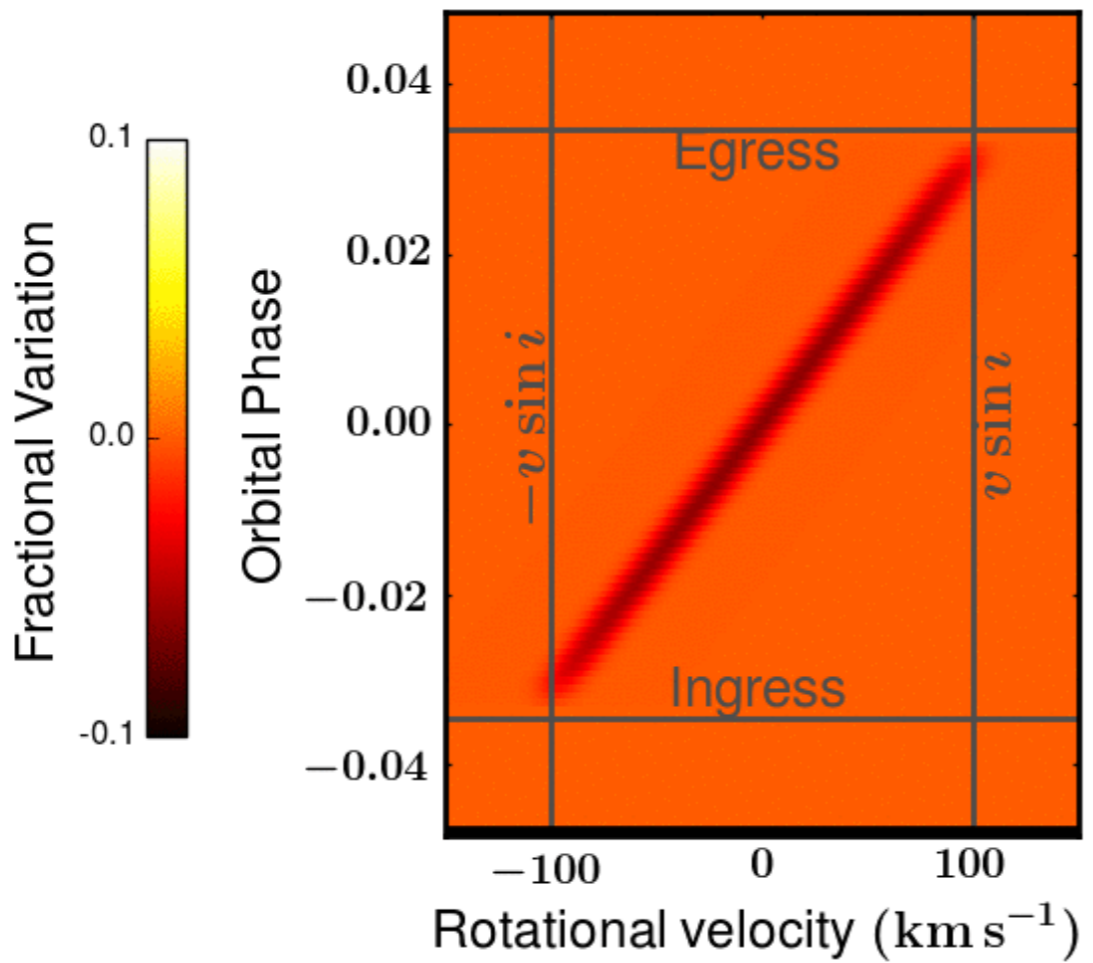
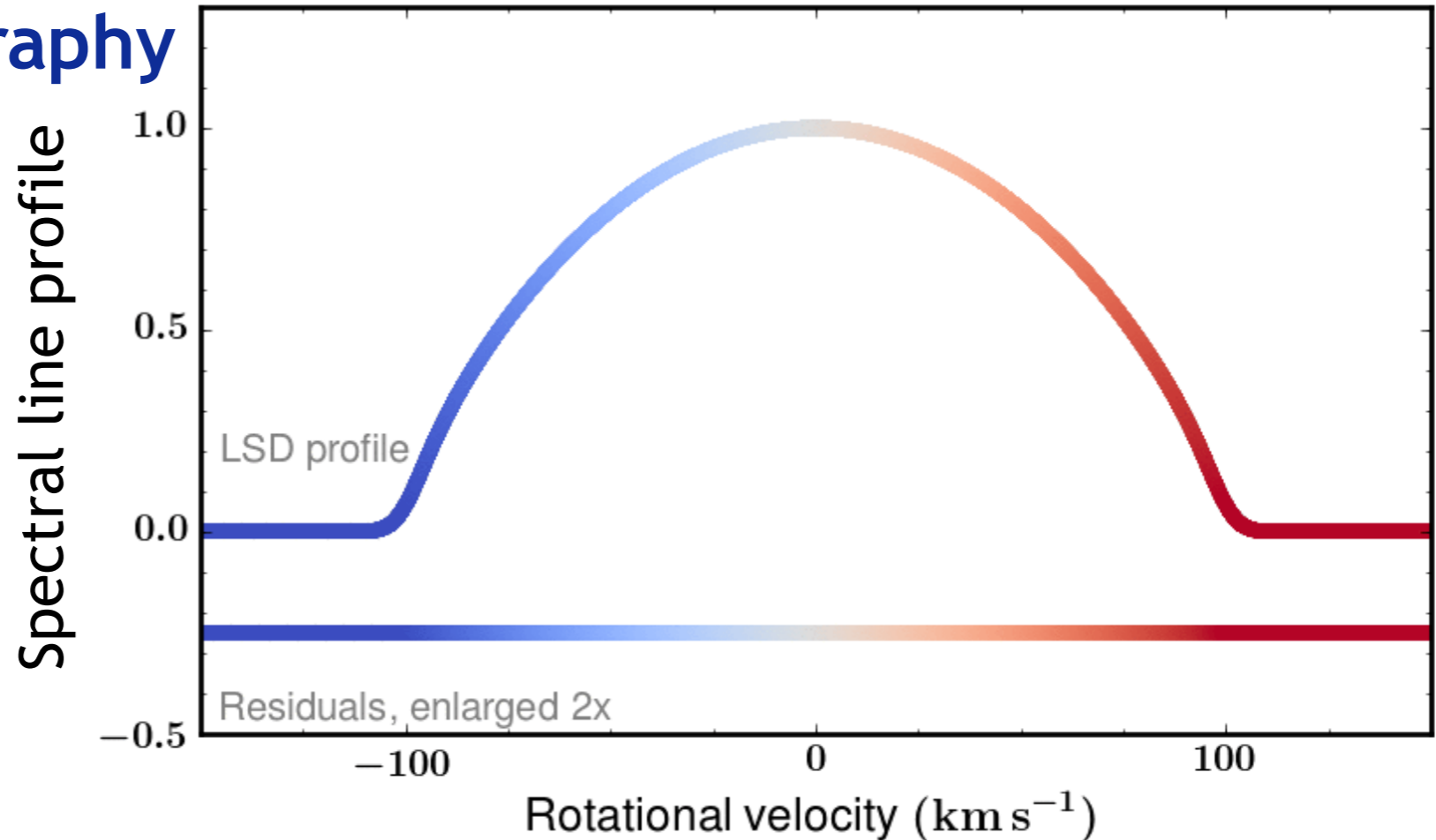




Early type stars are rapidly rotating

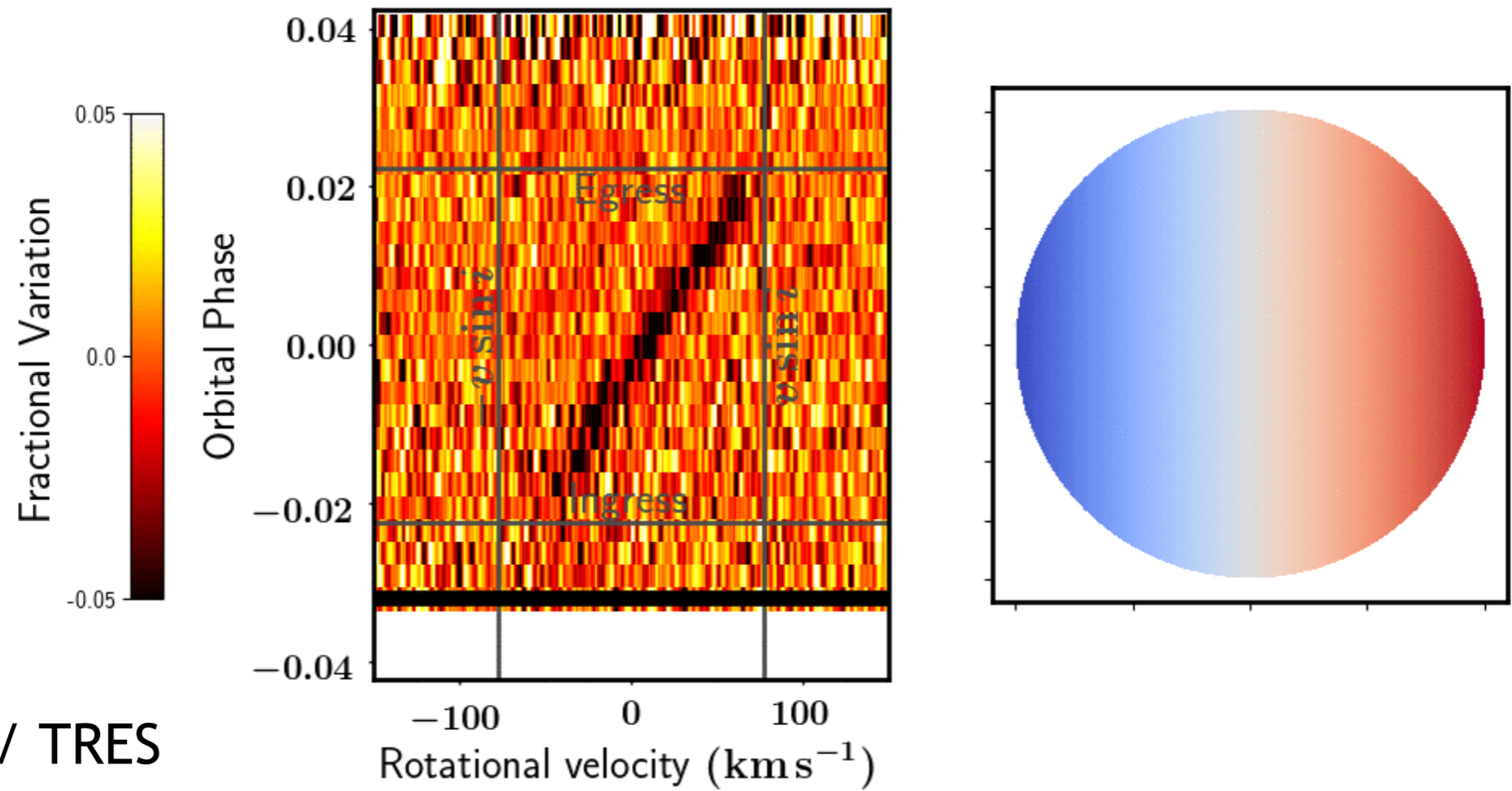
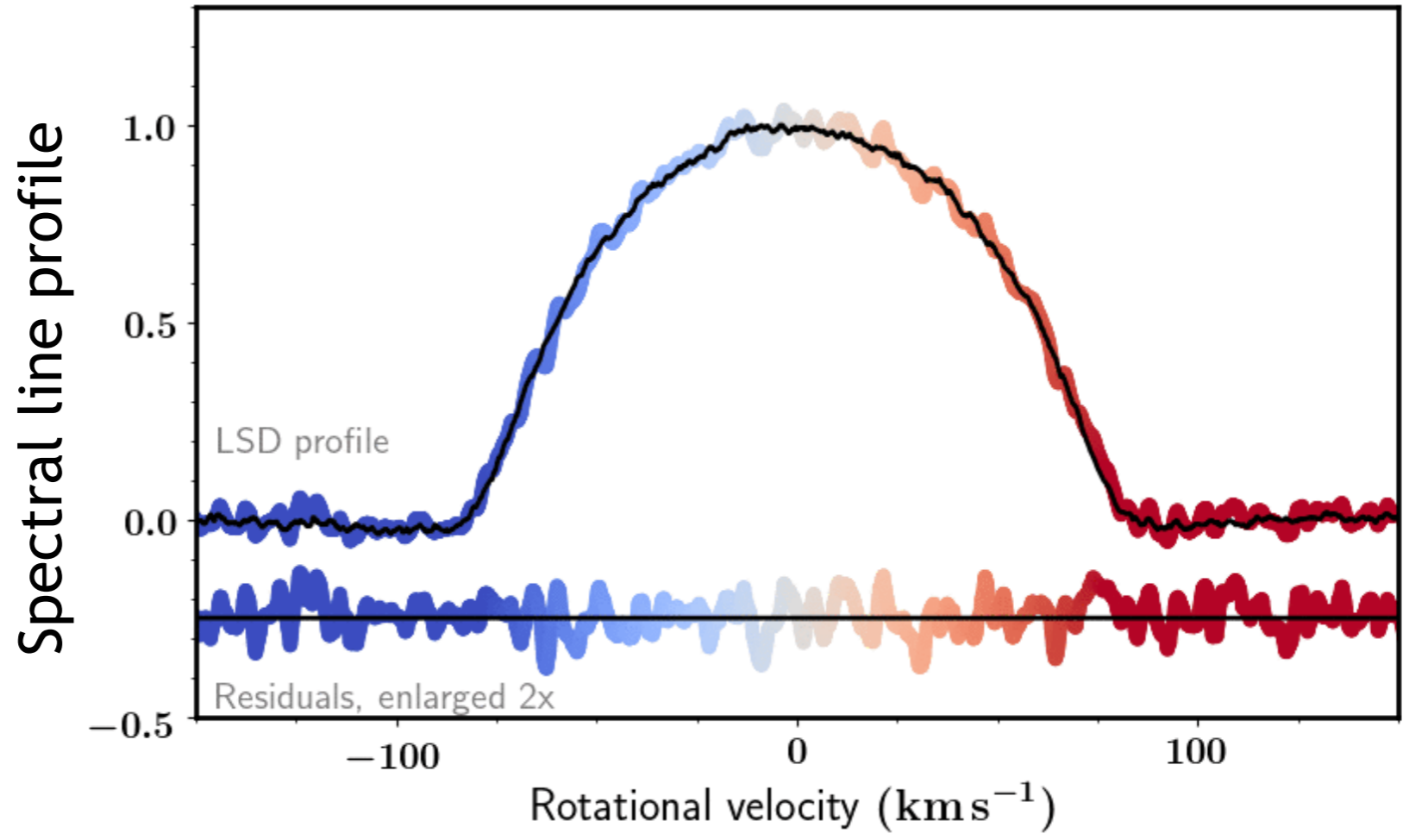


Doppler tomography



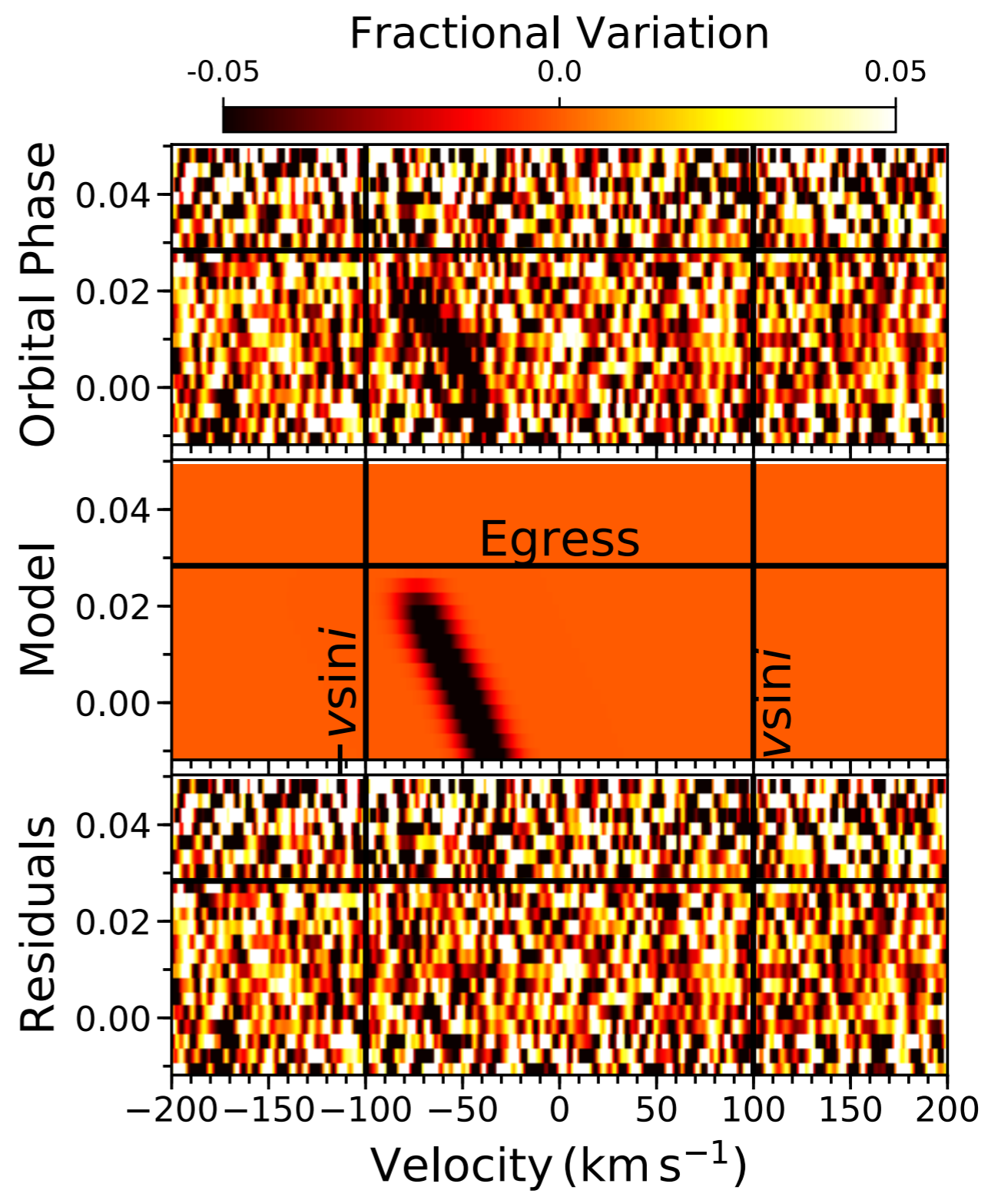
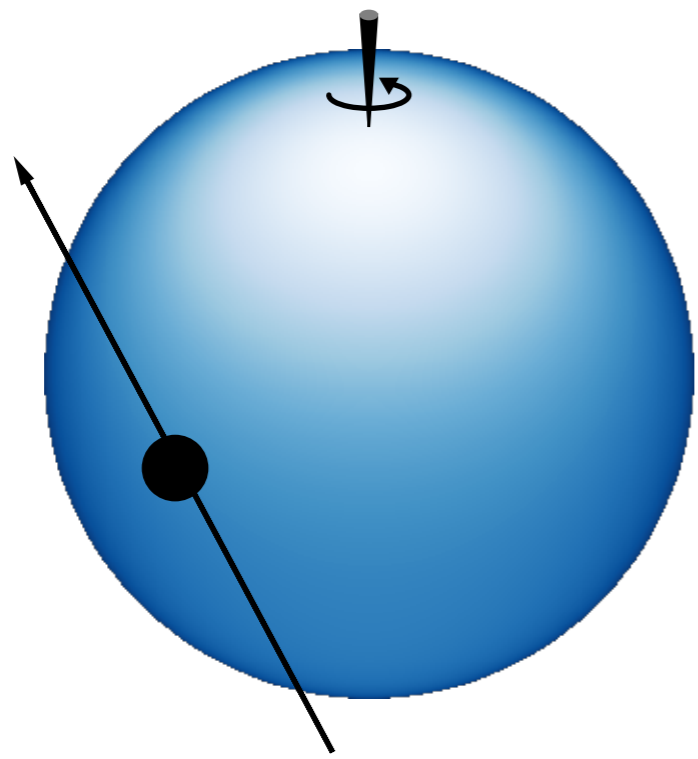
See Donati+ 1997
Collier Cameron+ 2010

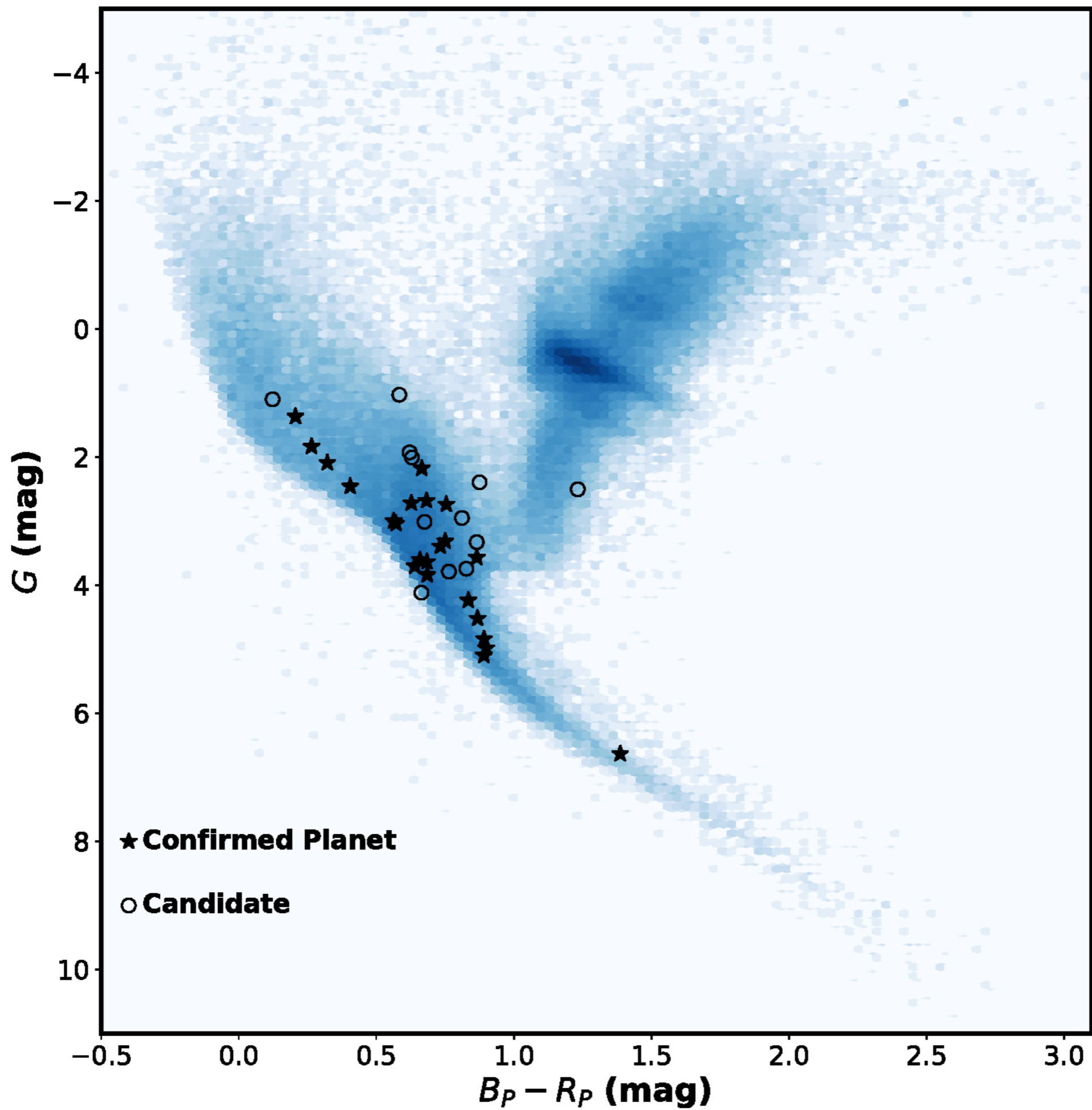
TOI-625 HAT-P-69b

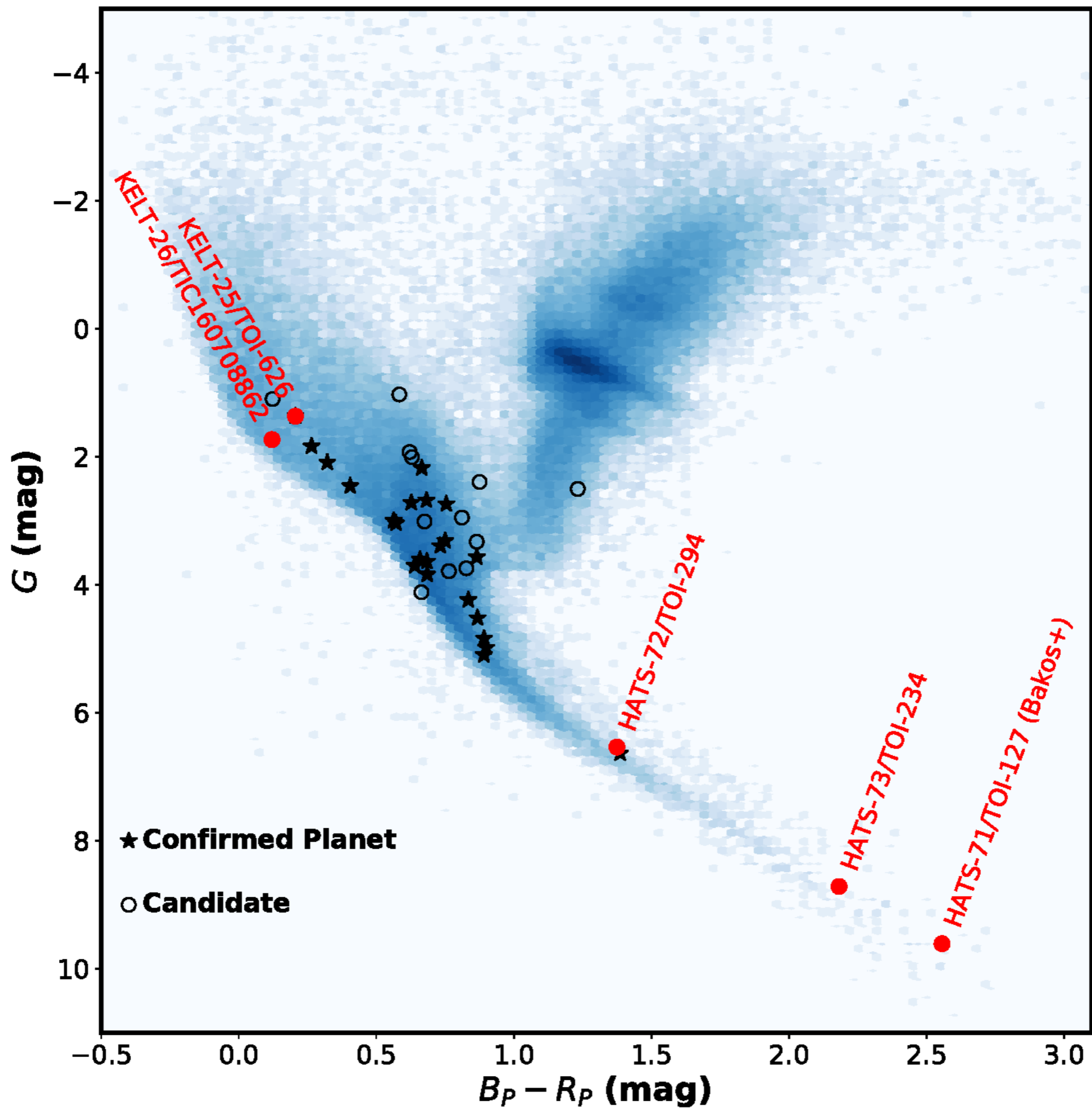


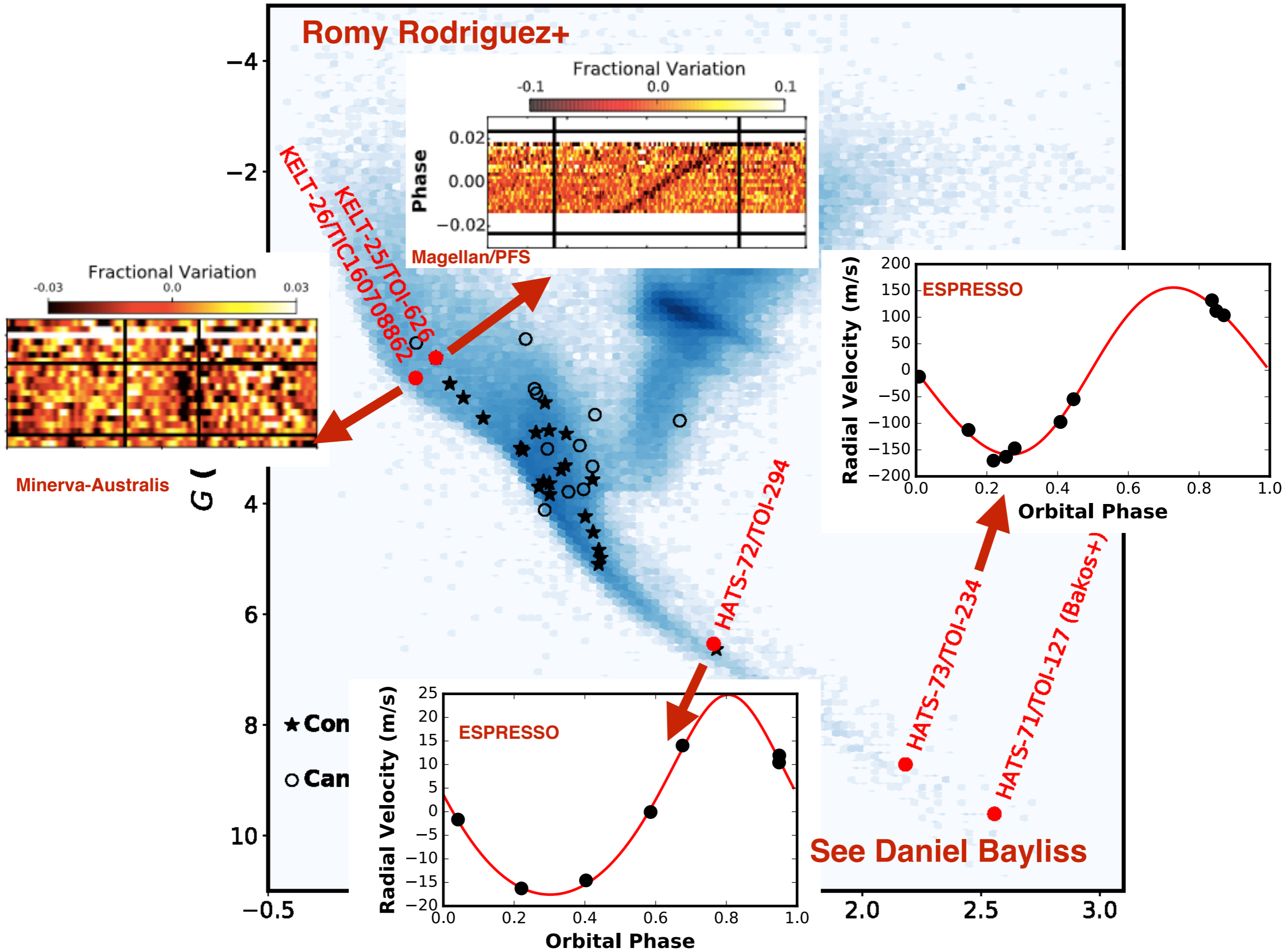
1.5m FLWO / TRES

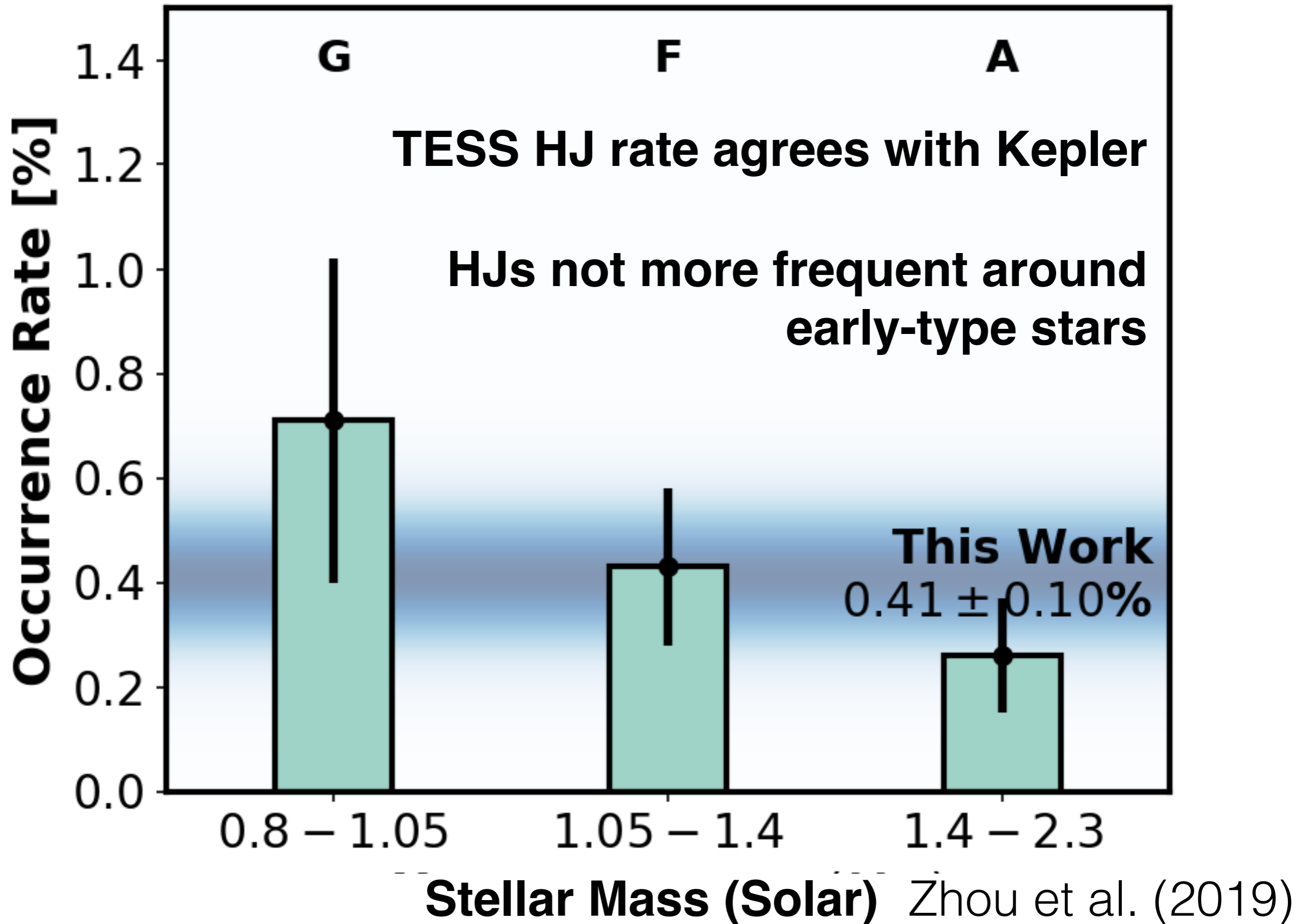
TOI-624 HAT-P-70b

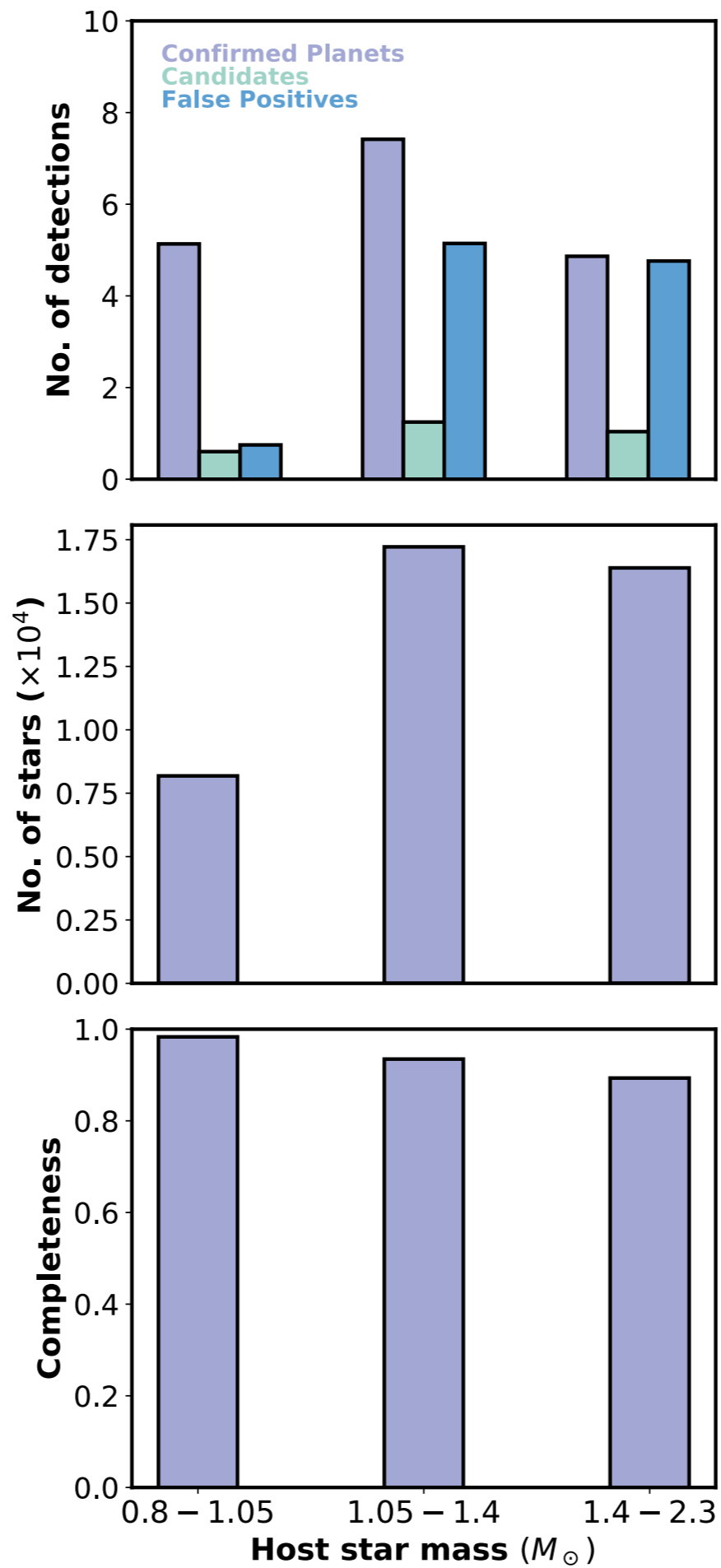


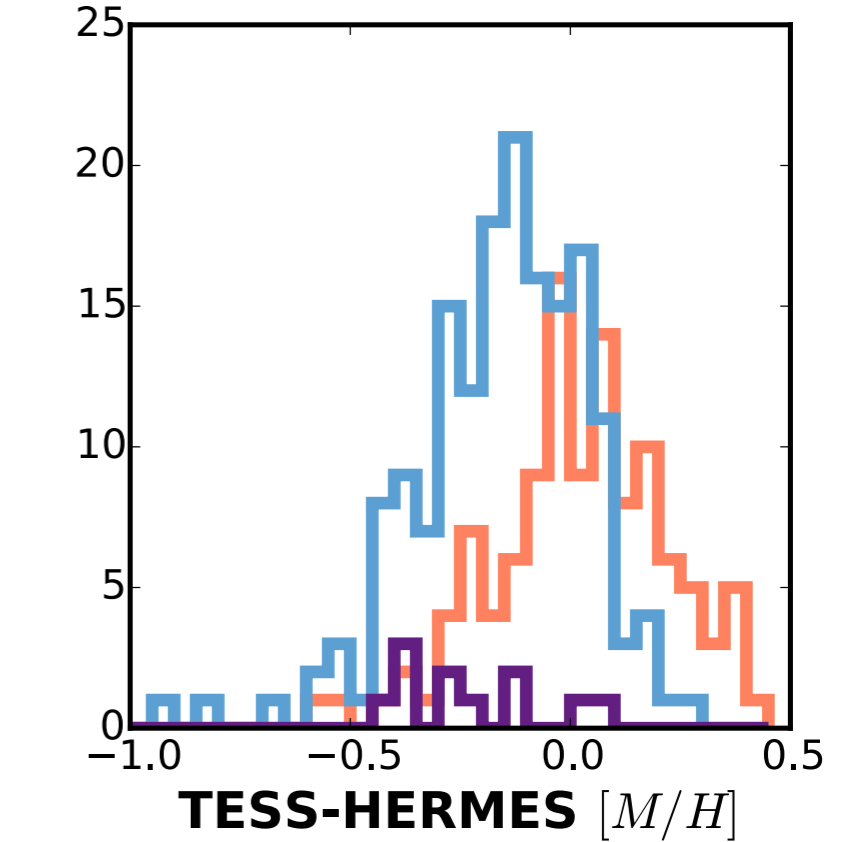
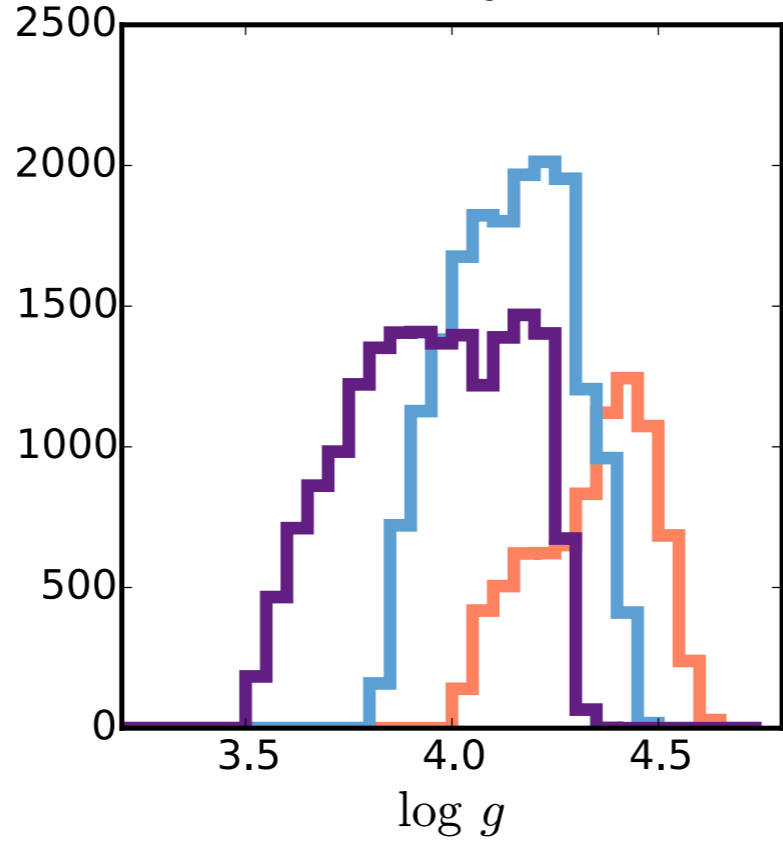
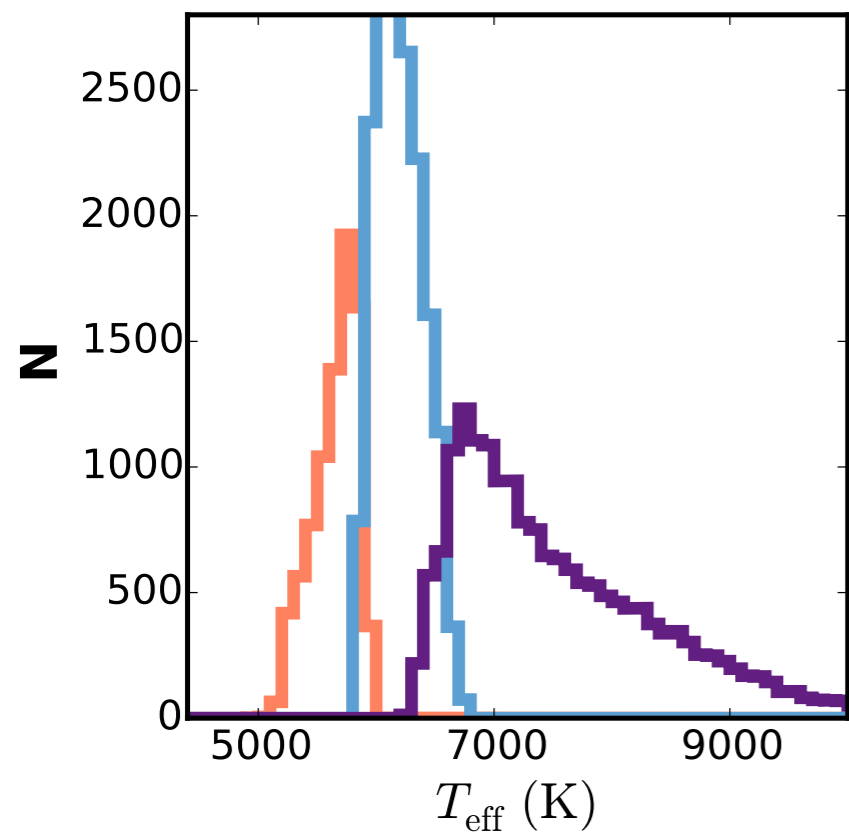
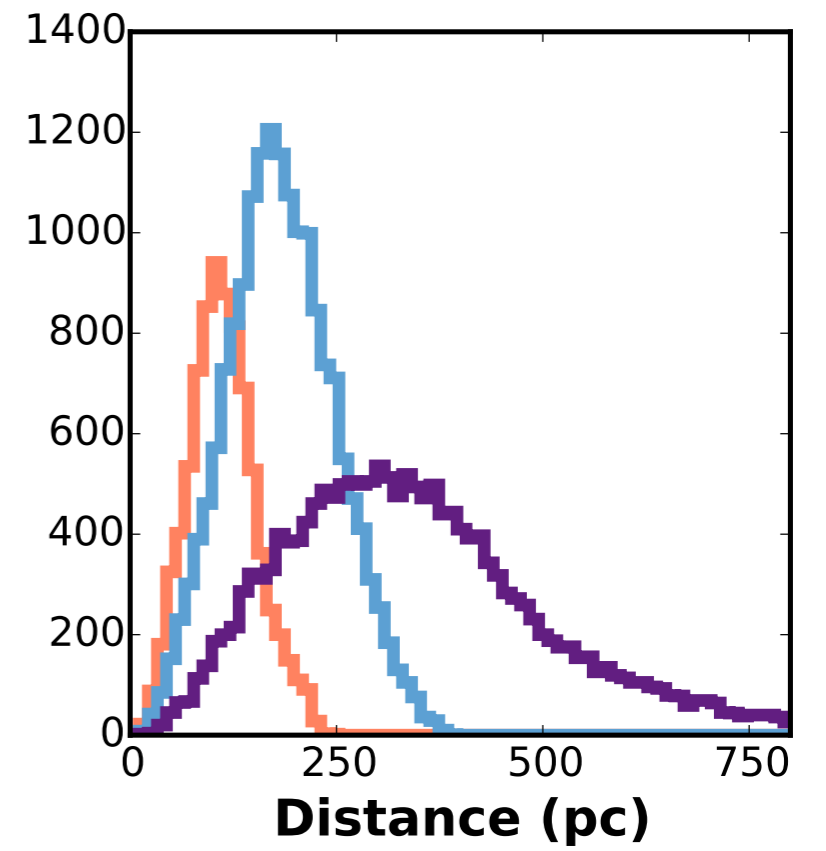
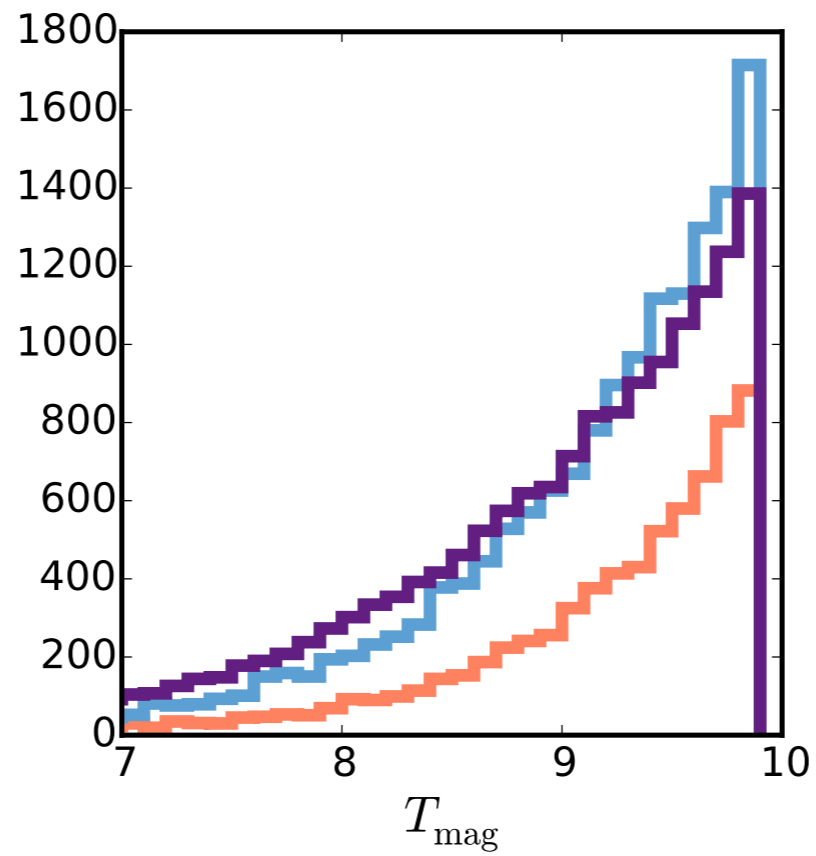
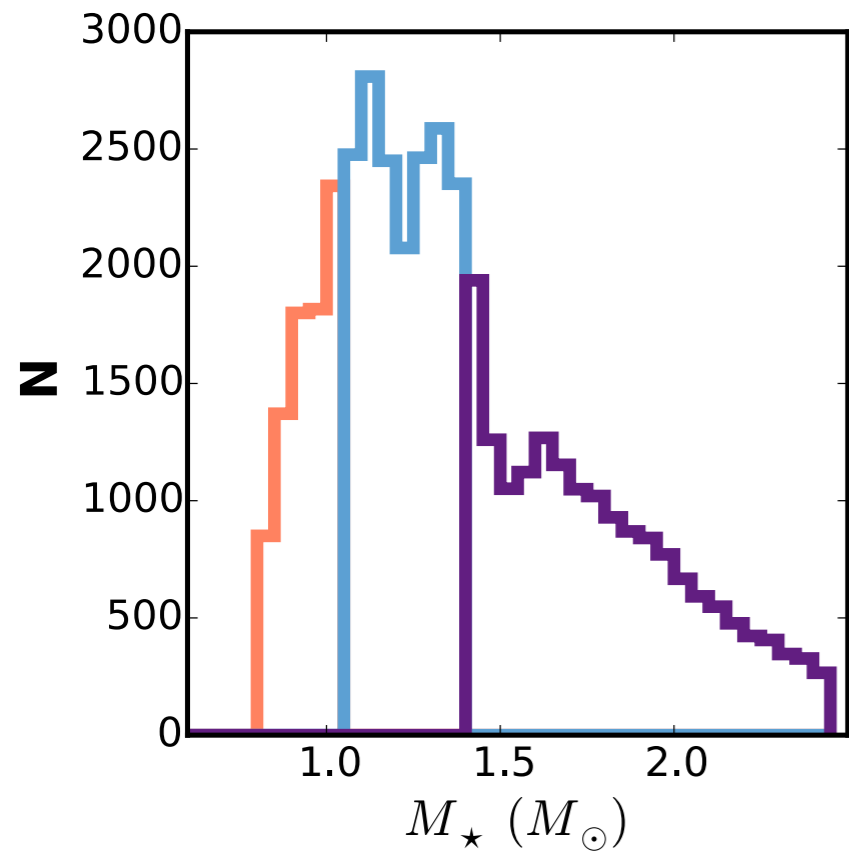


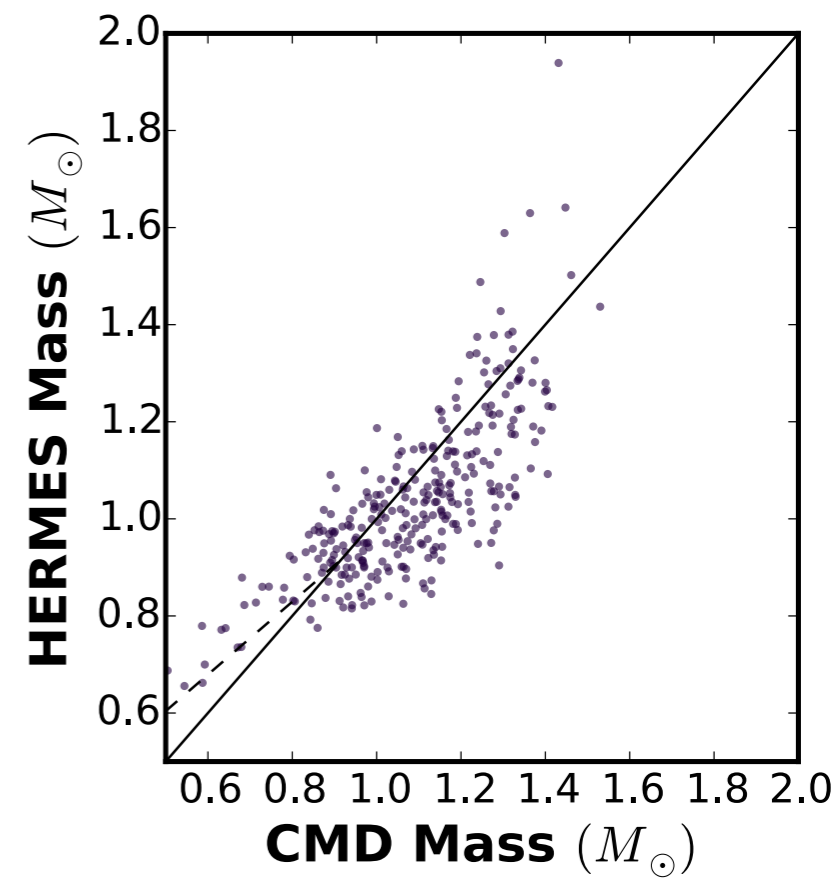
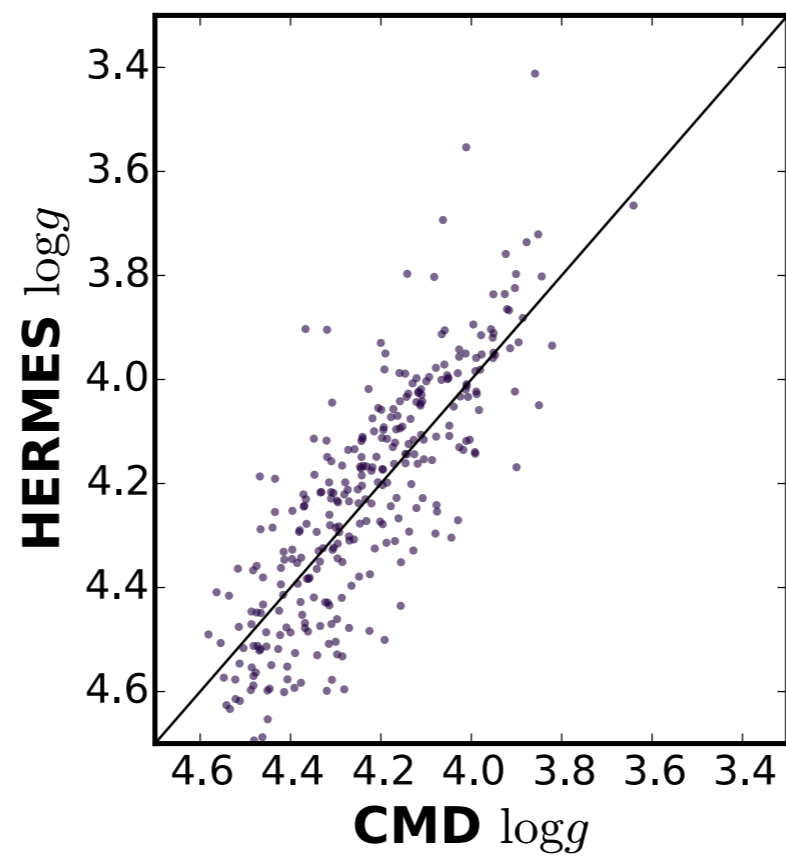
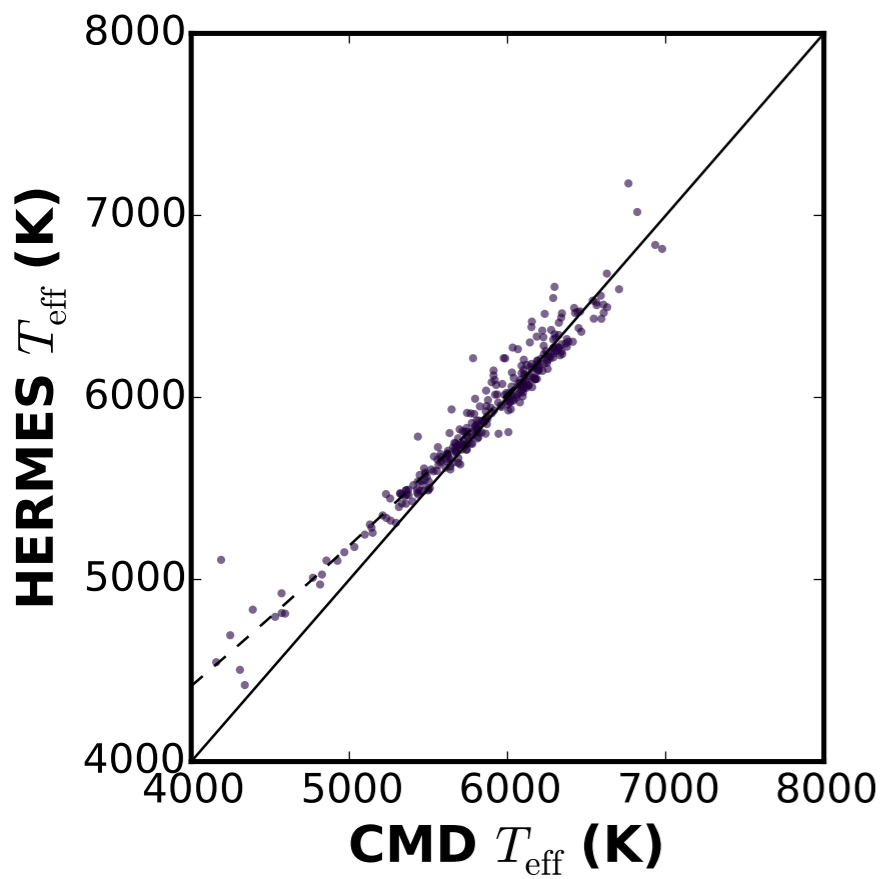






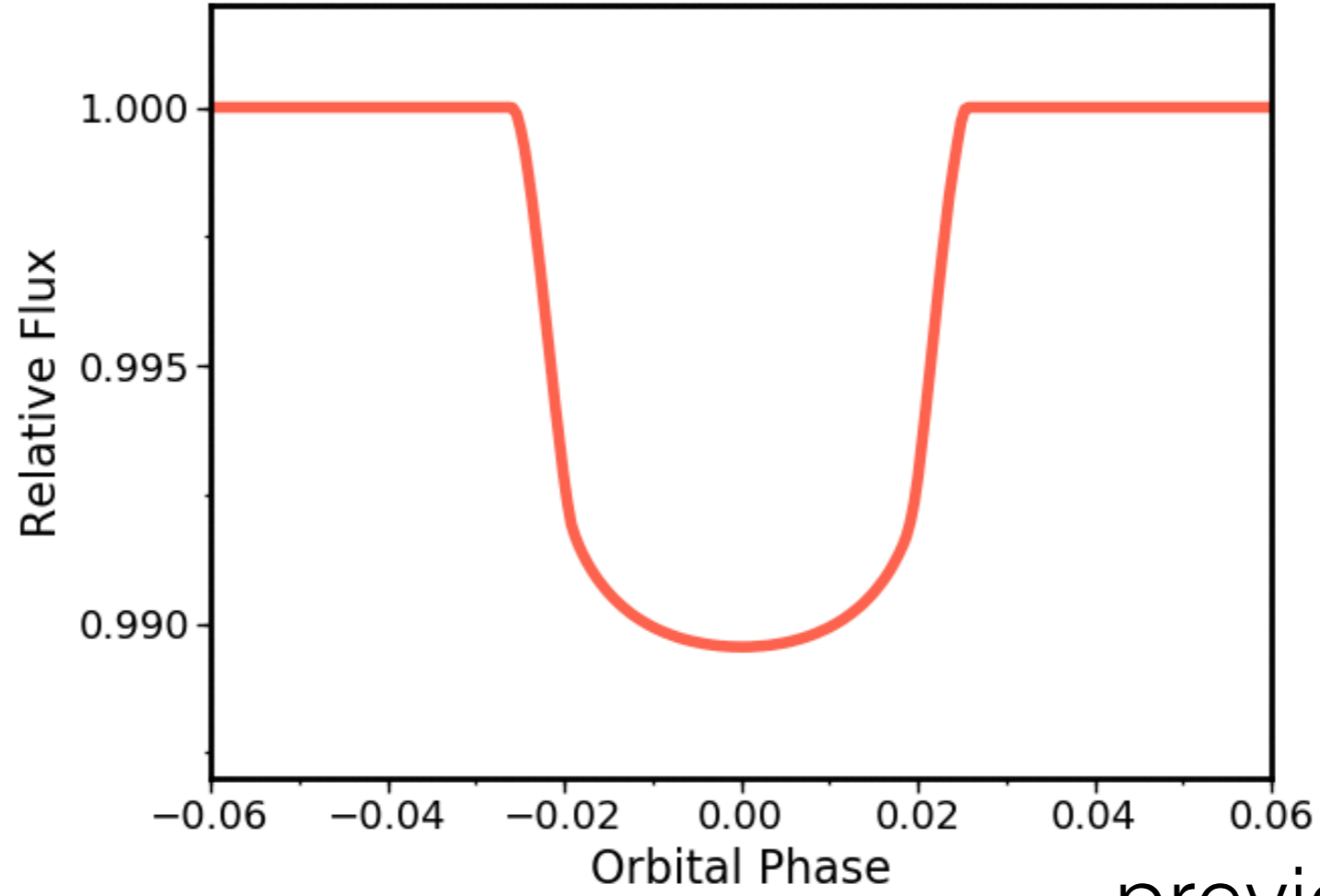
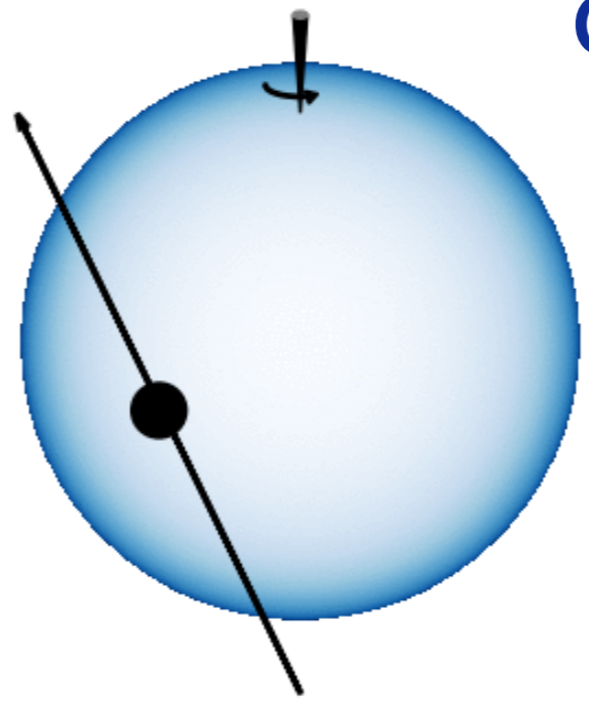






TOI-624
HAT-P-70b

Gravity darkening effect



See Barnes+ 2009

previously detected
for Kepler-13

TOI-624 HAT-P-70b

